Production rate calculations for cosmic-ray-muon-produced ¹⁰Be and ²⁶Al benchmarked against geological calibration data

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Abstract

First, I benchmark existing methods of calculating subsurface ²⁶Al, ¹⁰Be, and ¹⁴C production rates due to cosmic-ray muons against published calibration data from bedrock cores and mine excavations. This shows that methods based on downward propagation of the surface muon energy spectrum fit calibration data adequately. Of these methods, one that uses a simpler geographic scaling method based on energy-dependent attenuation of muons in the atmosphere appears to fit calibration data better than a more complicated one that uses the results of a global particle transport model to estimate geographic variation in the surface muon energy spectrum. Second, I show that although highly simplified and computationally much cheaper exponential function approximations for subsurface production rates are not globally adequate for accurate production rate estimates at arbitrary location and depth, they can be used with acceptable accuracy for many exposure-dating and erosion-rate-estimation applications.

Keywords: cosmogenic-nuclide geochemistry, exposure-age dating, erosion rate measurement, production rate calibration, beryllium-10, aluminum-26, carbon-14

1. Introduction

Naturally occurring cosmic-ray-produced nuclides that are useful for geochronology and other Earth science applications, for example ²⁶Al and ¹⁰Be, are produced in part by interactions with cosmic-ray muons. At the Earth's surface, production of these nuclides is predominantly due to interaction with high-energy neutrons, and muon interactions account for a small fraction of total production. However, the stopping distance of muons in rock is much longer than that of neutrons, so in the subsurface below a few meters depth, muon interactions account for nearly all cosmogenic-nuclide production. Thus, for geological applications of cosmogenic-nuclide geochemistry in which samples are exposed to the subsurface cosmic-ray flux, it is necessary to accurately estimate production rates due to muon interactions.

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Several different methods for computing production rates due to muon interactions are widely used at present. All these methods include calibrated parameters, either formally defined nuclear interaction cross-sections or empirical scaling constants based on simplified physics, that can variously be measured directly by muon irradiation of mineral targets in the laboratory or estimated from geological calibration data. Here 'geological calibration data' means measurements of concentrations of naturally occurring cosmic-ray-produced nuclides in settings where independent knowledge of the geological history of the site allows one to infer nuclide production rates from the concentration measurements.

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The purpose of this paper is to benchmark currently available muon production rate calculation methods against the existing set of geological calibration data, as a means of determining (i) how accurate the calculation methods are, and (ii) whether they are accurate enough for common geological and geochronological applications. First, I fit each production rate calculation method to calibration data in order to determine whether or not the method adequately fits the observations, and also to estimate values of interaction cross-sections or other fitted parameters. Second, based on these results, I evaluate the suitability of each method for common applications.

In this paper I consider production by muon interactions of ¹⁰Be, ²⁶Al, and ¹⁴C in quartz. Although there exist cross-section estimates and geological calibration data that can be used to also estimate cosmic-ray muon production of ³⁶Cl and ²¹Ne in various target minerals, I do not discuss these nuclides.

2. Muon production rate calculation methods.

Muon production of ¹⁰Be, ²⁶Al, and ¹⁴C in quartz takes place via two processes: capture of stopped negative muons (henceforth, "negative muon capture") and reactions with high-energy muons (henceforth, "fast muon interactions"). Dunai (2010) provides a comprehensive summary and Heisinger et al. (2002a,b) a detailed description of the production mechanisms; a brief summary follows. As muons originally incident on the Earth's surface travel through rock, they gradually lose energy by a variety of nuclear interaction processes and eventually come to a stop. At this point negative muons may be captured by a target nucleus to produce nuclides of interest. "Fast" muons, that is, muons that have not yet stopped of their own accord due to energy loss, may also produce these nuclides by a variety of higher-energy reactions. Muons incident at the Earth's surface span both a range of incidence angles and a wide spectrum of energies from near zero to very high energies capable of penetrating to depths of thousands of meters. Thus, both negative muon capture and fast muon interactions occur at all depths. Muons with lower energies and/or flatter incidence angles will stop at shallower depths, so as the depth below the surface increases, the remaining muon flux is more energetic and more collimated. For the nuclides of interest here, negative muon capture accounts for the majority of muon production in the upper few meters below the surface. However, the production rate due to negative muon capture decreases more rapidly with depth than fast muon production, so fast muon production is more important at greater depths.

Two classes of methods for calculating muon production rates are in use for Earth science applications at present. The first class is described by Heisinger et al. (2002a,b). 55

This approach (i) begins with a specified energy spectrum and angular incidence distribution of cosmic-ray muons at the Earth's surface, (ii) computes the stopping rate of muons with particular energy and incidence angle as a function of depth below the surface, and then (iii) integrates over energy and incidence angle to obtain integrated muon fluxes and stopping rates as a function of depth. One can then multiply the stopping rate at a given depth by a likelihood of nuclide production by negative muon capture, and multiply the muon flux at a given depth by a cross-section for nuclide production by interaction with fast muons, to obtain nuclide production rates by these two processes. This method has two important features. First, it accurately represents what is happening physically, specifically that the muon flux becomes more collimated and more energetic with depth. Second, the interaction cross-sections can be determined from laboratory experiments using artificially generated muon beams. Heisinger and others carried out these experiments, and these cross-section measurements are tabulated in the two papers cited above. This method, therefore, allows one to calculate production rates in natural settings entirely from the surface muon spectrum and parameters determined in laboratory experiments, so, in principle, no geological calibration is necessary (however, as discussed later, it appears that geological calibration provides more accurate cross-section estimates for the muon energy range of interest for geological applications). The key disadvantage of this method is that it is computationally somewhat time-consuming because computing the muon fluxes at a particular location and depth requires a numerical integration. For any application that requires integrating production rates during a long exposure history in which a sample's depth varies over time, computation time can be significant.

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In this paper I consider two implementations of the Heisinger method. Both use the same method of integration over energy and incidence angle, but they differ in how the muon spectra at the surface are specified.

One implementation, henceforth called "Method 1A," is a MATLAB implementation from Balco et al. (2008) in which the muon energy spectrum at the surface varies with atmospheric pressure exponentially with energy-dependent attenuation lengths taken from Boezio et al. (2000). In this method, the muon energy spectrum does not vary with position in the Earth's magnetic field. Specifically, in Method 1A I use the MATLAB scripts 'P_mu_total.m' and 'P_mu_total_alpha1.m,' version 1.2, dated September 2016 (all computer code used or described in this paper is available online; see Section 12).

The second implementation, henceforth "Method 1B," is a modification of this code by Lifton et al. (2014) in which the muon energy spectrum at the surface varies with atmospheric pressure, geomagnetic cutoff rigidity, and solar modulation according to Sato et al. (2008). Note that solar modulation is relatively unimportant for muon production, and in applying Model 1B throughout this paper I assume that the solar modulation constant is always the mean Holocene value according to Lifton et al. (2014). Specifically, Method 1B uses the MATLAB function 'P_mu_totalLSD.m,' version 1.0, dated March 2011, and an accompanying parameters file dated October 2013, both supplied by Nat Lifton.

Both Methods 1A and 1B use the same method for calculating the subsurface muon fluxes and stopping rates, and thence the nuclide production rates, from a specified surface muon energy spectrum. The difference between them, as described in the preced-

ing paragraphs, lies in how the muon energy spectrum at the surface is specified. Both implementations fully specify the muon flux and stopping rate at any arbitrary location and depth below the surface. To compute nuclide production rates, then, they require measured or calibrated values for (i) a likelihood of production by capture of stopped negative muons, and (ii) a cross-section for production by fast muon interactions, for the nuclide-target mineral pair of interest.

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The second class of calculation methods does not compute the muon flux by in-108 tegration of surface spectra as is done in the Heisinger method. In fact, it does not 109 compute muon fluxes or stopping rates at all. Instead, muon production rates are sim-110 ply approximated by an exponential function, or a sum of several exponential functions, 111 in mass depth below the Earth's surface. This approach has been used by several au-112 thors, for example, by Granger and Smith (2000) and, recently, Braucher et al. (2011) 113 and Braucher et al. (2013). The advantage of this method is that it is computationally 114 trivial, so lends itself to situations in which cosmogenic-nuclide production rates must 115 be integrated over time as a sample changes depth. The disadvantage is that it does not 116 correctly represent the physics of muon production: because the muon flux becomes 117 more collimated and more energetic with increasing depth, the instantaneous e-folding 118 length for muon production continually increases with depth. This effect cannot be 119 represented with a finite sum of exponential functions. Mainly, the importance of this 120 method is that it is computationally extremely simple and also, most likely, accurate 121 enough for many geological applications. 122

A recent implementation of the exponential-approximation approach (Braucher 123 et al., 2013) computes subsurface production rates due to muons for a particular nuclide-124 mineral pair by assuming that total muon production as a function of depth at a spe-125 cific site is exponential in mass depth and defined by a surface production rate and 126 a subsurface e-folding length. They found that apparent subsurface e-folding lengths 127 were similar at various locations where subsurface calibration data existed (with an 128 average of 4380 \pm 650 g cm⁻² for ¹⁰Be and ²⁶Al, excluding one outlier), and they 129 assumed that the surface production rate varies exponentially with atmospheric depth 130 with an e-folding length in the atmosphere obtained from previously published atmo-131 spheric muon flux measurements. Henceforth I will refer to this approach as 'Method 132 2.' Method 2 is purely empirical and defines subsurface production rates at arbitrary 133 location and depth for a particular nuclide-mineral pair as a function of two parameters 134 derived from fitting to geological calibration data: a reference surface production rate 135 at 1013.25 hPa atmospheric pressure, and a subsurface e-folding length. Compared to 136 other exponential-approximation methods, this method has two advantages. One, it is 137 very simple. Two, in contrast to most other published applications of the exponential-138 approximation method, it includes a geographic scaling method. 139

3. Geoscience applications that require muon production rates to be calculated.

Here I consider three classes of Earth science applications of cosmogenic-nuclide geochemistry that require calculations of production rates due to muons: surface exposure dating; surface erosion rate measurements; and burial dating (with the related application of depth-profile dating). These classes of applications have very different 142 requirements for how precise the muon production rate calculations must be to maintain the desired precision in the eventual quantity being measured, e.g., an age or an erosion rate. Potentially, therefore, some applications could accept reduced precision in calculating production rates due to muons in exchange for computational speed and simplicity, without significantly compromising the overall results.

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Exposure-dating applications are for nearly all practical purposes concerned with 150 surfaces that have low erosion rates. Thus, the cosmogenic-nuclide concentration that 151 is measured is the result of production very close to the surface, where production is 152 nearly all by high-energy neutron spallation. At most a few percent of surface produc-153 tion is due to muon interactions (with the exception of in-situ-produced cosmogenic 154 14 C in quartz, where ca. 10% of surface production is due to muons). At present, typ-155 ical measurement uncertainty on ¹⁰Be concentrations in exposure-dating applications, 156 for example, is 3%, and typical uncertainties in estimating total surface production 157 rates are at least 5%, leading to total uncertainties in exposure age near 6% in most 158 cases. Suppose our goal is for uncertainty in muon production rates to contribute only 159 10% of this total uncertainty. If muon production is 2% of total surface production, 160 we can accept a 30% uncertainty in estimating the surface muon production rate and 161 still achieve this goal. Thus, exposure-dating applications do not require very precise 162 estimates of surface production rates due to muons, and do not require estimates of 163 subsurface production rates at all. 164

Erosion-rate applications exploit the fact that surface erosion advects subsurface 165 rock or soil through a near-surface layer in which cosmogenic-nuclide production oc-166 curs. The faster the erosion rate, the shorter its residence time in the production zone, 167 and the lower the nuclide concentration in a particular package of material when it 168 reaches the surface. Thus, the surface nuclide concentration is inversely proportional 169 to the erosion rate. Relating surface nuclide concentrations to erosion rates requires 170 computing the integrated production during the entire subsurface residence time of the 171 sample, a significant part of which is typically spent below several meters depth where 172 production is entirely by muons. At high erosion rates, up to $\sim 25\%$ of the total surface 173 nuclide concentration can be muon-produced (note that this calculation applies to ¹⁰Be 174 and ²⁶Al: this fraction would be larger for *in-situ*-produced ¹⁴C, although this nuclide 175 is not generally used for erosion rate estimates). Thus, in relation to the case of surface 176 exposure dating where all relevant production takes place at the surface, maintaining 177 desired precision on an erosion rate estimate requires better precision on estimating 178 muon production. However, we are not concerned with how precise the estimate of the 179 muon production rate is at any particular depth, but rather with how precisely we have 180 computed the integrated muon production over the entire period in which the sample 181 is advected to the surface. Thus, we could potentially accept a method that was inac-182 curate at any particular depth, but resulted in an accurate computation of the integral 183 production. 184

Burial-dating applications (Granger, 2006), and also the related category of depthprofile dating applications (see summary in Hidy et al., 2010) are the most demanding application from the perspective of computing production rates due to muons. This is mainly because they involve subsurface samples whose measured nuclide concentration may be nearly entirely muon-produced. In addition, in many of this class of applications, samples have resided at the same depth for long periods of time, so it is necessary to accurately compute the production rate at a specific depth, rather than an 191 integrated production rate over a range of depths. Commonly in this application, the uncertainty in the burial (or depth-profile) age might be dominated by the uncertainty 193 in the estimate of subsurface production rates by muons: in contrast to the situation of surface exposure dating discussed above, a 30% uncertainty in estimating production 195 rates due to muons might contribute a 10-15% uncertainty to a burial age or a 30% uncertainty to a depth-profile age. To summarize, although there exist some variants 197 of burial dating that are less sensitive to this issue (e.g., the isochron method of Balco 198 and Rovey, 2008), these applications are, in general, the most demanding from the per-199 spective of muon production rate calculations in that they require a precise estimate of 200 the muon production rate at a particular depth and location.

4. Geological calibration data.

4.1. What are geological calibration data exactly?

Geological calibration data that can be used to estimate production rates due to 204 muons consist of measured nuclide concentrations in rock that is approximately 1000 205 g cm⁻² or more below the surface, such that the measured concentration is entirely 206 due to muon production and not to high-energy neutron spallation. Ideally, one could 207 obtain calibration data from a site where the rock mass had experienced a single period 208 of exposure, of independently known duration, at the beginning of which the nuclide 209 concentration was zero, and during which zero surface erosion took place. So far it 210 has not been feasible to find a site that has all these properties and also where the 211 period of exposure has been long enough that nuclide concentrations below several 212 meters depth would be accurately measurable. Instead, therefore, existing calibration 213 data come from drill cores and mine excavations at sites where (i) the surface erosion 214 rate is relatively low, and (ii) geologic evidence indicates that the surface has been 215 steadily eroding at this low rate for an extended time, ideally millions of years, such 216 that nuclide concentrations have reached a steady state in which nuclide production 217 at a certain depth is balanced by loss by radioactive decay and advection of material 218 towards the surface due to erosion. 219

In this case, the workflow goes as follows. First, measure the surface nuclide con-220 centration, which is nearly all due to spallogenic production, and apply an indepen-221 dently calibrated spallogenic production rate estimate to determine the erosion rate. 222 Second, measure the nuclide concentration in the subsurface where spallogenic pro-223 duction is zero. The nuclide concentration is related to the production rate as follows: 224

$$N_i(z) = \int_0^\infty P_i(z + \epsilon t) e^{-\lambda_i t} dt$$
(1)

Here $N_i(z)$ is the measured concentration of nuclide *i* (atoms g⁻¹) at depth *z* (g 225 cm⁻²), $P_i(z)$ is the production rate of nuclide *i* (atoms g⁻¹ yr⁻¹) at depth *z*, λ_i is the 226 decay constant of nuclide i (yr⁻¹), ϵ is the erosion rate (g cm⁻² yr⁻¹), and t is time (yr), 227 which is zero at the present time and positive for past times. If we have a prediction 228 for the function $P_i(z)$, we can evaluate it by using this relationship to calculate the 229 expected nuclide concentration at the sample depth and comparing it to the measured 230

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concentration. In simplified cases this can be solved directly for $P_i(z)$. For example, if $P_i(z)$ is assumed to be locally exponential in *z* with an e-folding length Λ , such that $P_i(z + \delta z) = P_i(z) \exp(-\delta z/\Lambda)$, then $P_i(z) = N_i(z)(\lambda_i + \epsilon/\Lambda)$ (also see discussion below). Alternatively, if the erosion rate is zero, then $P_i(z) = N_i(z)\lambda_i$ no matter what the form of $P_i(z)$.

The main difficulty with this approach is that it requires the assumption that the 236 erosion rate has been steady for long enough for nuclide concentrations to reach a 237 steady state in which nuclide production at given depth is balanced by radioactive de-238 cay and advection of material towards the surface by erosion. For ¹⁰Be and erosion 239 rates on the order of meters per million years, for example, this implies steady ero-240 sion for several million years, which is probably unlikely in nearly all real geological 241 situations. As a practical matter, it is impossible to verify this assumption. However, 242 one can mitigate the dependence on this assumption by looking for calibration sites 243 with the lowest possible erosion rate. For one thing, this is useful because the lower 244 the erosion rate, the higher the steady-state nuclide concentration, and thus the easier 245 and the more precise the measurement. This is particularly important from the per-246 spective of estimating subsurface production rates, because nuclide concentrations at 247 several meters depth are orders of magnitude below those in surface samples. More 248 importantly, the lower the erosion rate, the less sensitive the production rate estimate 249 is to both the assumed value of the erosion rate and the assumption of steady erosion. 250 A heuristic explanation for this is just that at very low erosion rates, nuclide concen-251 trations are close to equilibrium between production and radioactive decay (the rate of 252 which is known accurately). Thus, loss by erosion is a relatively small fraction of the 253 nuclide balance, so uncertainty in the magnitude of loss by erosion has a commensu-254 rately small effect on the precision of the total production rate estimate. To show this 255 quantitatively, simplify the production rate-depth relationship $P_i(z)$ as discussed above 256 such that $P_i(z + \delta z) = P_i(z) \exp(-\delta z/\Lambda)$, where Λ is an e-folding length (g cm⁻²). In 257 this case (see Lal, 1991), Equation 1 is: 258

$$N_i(z) = \frac{P_i(z)}{\lambda_i + \frac{\epsilon}{\Lambda}}$$
(2)

Thus, one can compute $P_i(z)$ from $N_i(z)$ by:

$$P_i(z) = N_i(z)\lambda_i + \frac{\epsilon N_i(z)}{\Lambda}$$
(3)

The (absolute) uncertainty in P_i , σ_P , that is derived from the (absolute) uncertainty in the erosion rate measurement σ_{ϵ} is:

$$\sigma_P = \sigma_\epsilon \frac{N_i(z)}{\Lambda} \tag{4}$$

where σ_{ϵ} is the uncertainty (g cm⁻² yr⁻¹; here uncertainty is assumed to be Gaussian) in the erosion rate estimate. Given $\Lambda = 1500$ g cm⁻² (appropriate for production by negative muon capture near ~1000 g cm⁻² depth) and $\Lambda = 4200$ g cm⁻² (appropriate for production by fast muons at greater depth), Figure 1 shows the relative uncertainty in the production rate estimate (that is, σ_P/P) resulting from a 50% uncertainty in the erosion rate estimate (that is, $\sigma_{\epsilon} = 0.5\epsilon$). At low erosion rates where nuclide loss by

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Figure 1: Relative uncertainty in production rate estimate (σ_P/P) for ¹⁰Be resulting from a 50% uncertainty (that is, a relative uncertainty of 0.5) in the erosion rate estimate, for two different local e-folding lengths Λ (see text). $\Lambda = 1500 \text{ g cm}^{-2}$ approximates negative muon capture in the shallow subsurface (order 1000 g cm⁻² depth), whereas 4200 g cm⁻² approximates fast muon production at greater depths. In addition, the black line shows the equivalent result for ¹⁴C with $\Lambda = 1500 \text{ g cm}^{-2}$. Names of calibration sites discussed later in the text indicate erosion rates characteristic of each site. Rock density assumed 2.7 g cm³.

radioactive decay is an important part of the nuclide balance, the uncertainty in esti-268 mating P_i is strongly dependent on the erosion rate, so lower erosion rate sites incur 269 significantly lower uncertainties in production rate estimates. The uncertainty in esti-270 mating P_i also decreases for larger Λ , and the instantaneous value of Λ increases with 271 increasing depth below the surface, so, for a given erosion rate, uncertainties in esti-272 mating the production rate will also decrease with increasing depth. The point of this 273 exercise is that it shows that even if we relax our steady-state assumption to only the 274 relatively weak assumption that the erosion rate has varied by less than 50% during 275 the time period in which the nuclide concentrations accumulated, if we can find a site 276 where the erosion rate is less than ca. 1 m Myr⁻¹, we can still estimate P_i with better 277 than ca. 10% precision. If we can find a site where the erosion rate is less than ca. 10 278 cm Myr⁻¹, we can estimate production rates with percent-level accuracy. In addition, 279 as shown in Figure 1, for a particular depth and erosion rate, uncertainty in estimat-280 ing the production rate is smaller for nuclides with shorter half-lives. Thus, one could 281 achieve desired precision in production rate estimates at higher erosion rates for ²⁶Al, 282 ³⁶Cl, and particularly ¹⁴C, than for ¹⁰Be. To summarize, however, for the most accurate 283 estimate of production rates due to muons, with minimal reliance on a strong steady-284 state assumption, we want calibration sites with the lowest possible erosion rates. One 285 could also obtain the same effect by choosing deeper samples at a site with a higher 286 erosion rate (I discuss this more later), but of course this would only permit accurate 287 production estimates for some nuclides at large depths. 288

4.2. Existing calibration data sets

I will discuss published calibration data that are relevant to estimating muon production rates of ¹⁰Be and ²⁶Al from five sites, as follows: 289 290

Beacon Heights, Antarctica. At this site, John Stone and colleagues collected a 292 25-m core in sandstone bedrock at 2183 m elevation in the McMurdo Dry Valleys of 293 Antarctica. ¹⁰Be and ²⁶Al measurements from this core by several laboratories are 294 reported in Borchers et al. (2016). Surface nuclide concentrations at this site indicate 295 an erosion rate near 0.1 m Myr⁻¹. This is by far the lowest erosion rate site for which 296 calibration data exist. Balco et al. (2011b) reported ²¹Ne measurements from this core, 297 but I do not discuss them here.

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La Ciotat, France. This site is a 10-m core in quartzose limestone bedrock at 310 m elevation. ¹⁰Be and ²⁶Al (and also ³⁶Cl, which is not discussed further here) concentrations in this core are described by Braucher et al. (2011). Surface nuclide concentrations at this site indicate an erosion rate near 40 m Myr⁻¹.

Leymon quarry, Spain. This site consists of two 25-m cores (the "Leymon Low" and "Leymon High" cores at 1246 and 1277 m elevation, respectively) drilled in a quartzite dike. Braucher et al. (2013) report ²⁶Al and ¹⁰Be measurements for both cores, and Lupker et al. (2015) report *in-situ*-produced ¹⁴C measurements for the Leymon High core. Surface nuclide concentrations here indicate erosion rates near 20 m Myr^{-1} .

Macraes Flat, New Zealand. This site is a series of mine excavations at 535 m 309 elevation. Kim and Englert (2004) reported ¹⁰Be and ²⁶Al measurements from samples up to ca. 180 m below the surface at this site. Although these authors concluded 311 that near-surface nuclide concentrations were not in steady state with a constant erosion rate, subsurface nuclide concentrations indicate an erosion rate near 10 m Myr⁻¹. 313 Although Kim et al. (2007) reported ¹⁴C concentrations in quartz at this site, they con-314 cluded that a significant fraction of their observed concentrations was due to thermal 315 neutron capture on N rather than muon interactions. Thus, I have not considered these 316 measurements further.

Cuiaba, Brazil. This site consists of two open-pit excavations at 210 m elevation. ¹⁰Be measurements from quartz veins at these sites are described by Braucher et al. (2003). Surface nuclide concentrations here indicate erosion rates on the order of 0.5-1 $m Mvr^{-1}$.

Figures 2 and 3 show ¹⁰Be and ²⁶Al data from these sites. I accept all mass depths 322 of samples as reported in the source papers, with the exception that I independently 323 carried out a depth adjustment for some samples at Cuiaba by fitting to a subsurface 324 production profile, as described in section 6.2.1. The nuclide concentration measure-325 ments have been renormalized to common ¹⁰Be and ²⁶Al measurement standards, as 326 follows. For ¹⁰Be, all measurements except those from Macraes Flat and Cuiaba were 327 originally measured against ¹⁰Be standards compatible with those of Nishiizumi et al. 328 (2007). Those from Macraes Flat were measured against the ¹⁰Be standards of Nishi-329 izumi (2002) (R. Finkel, written communication), so have been corrected by a factor 330 of 0.9042; those at Cuiaba were measured against the certified value of the NIST Be 331 isotope ratio standard, so have been corrected by a factor of 1.042 (Nishiizumi et al., 332 2007). ²⁶Al measurements in these studies employed both the standards of Nishiizumi 333 (2004) (measurements at LLNL-CAMS: all Macraes Flat and most Beacon Heights 334 data) and those of Arnold et al. (2010) (measurements at ASTER: La Ciotat, Leymon 335 Quarry, and some Beacon Heights data). ASTER ²⁶Al data have been normalized to 336 the KNSTD standardization using a correction factor of 1.021, which is based on com-337



Figure 2: ¹⁰Be calibration data from cores and excavations. All ¹⁰Be concentration measurements have been normalized to the '07KNSTD' standardization as described in the text, but have not been otherwise corrected or adjusted; variation in nuclide concentrations among sites reflects variation in site elevation and erosion rate. Note that the depths of samples in the deeper of two excavations at Cuiaba have been adjusted using the fitting procedure described later in the text. Error bars show 1- σ uncertainties as reported in the source publications; where not shown they are smaller than the symbols used to plot the data at this scale.



Figure 3: ²⁶Al calibration data from cores and excavations. All ²⁶Al measurements have been normalized to the 'KNSTD' standardization as described in the text. Error bars show $1-\sigma$ uncertainties as reported in the source publications; where not shown they are smaller than the symbols used to plot the data at this scale.

parison of the KNSTD Al standards with an Al standard derived from the same source material as the ASTER Al standards (Fink and Smith, 2007). Note that this renormalization of the ²⁶Al data has a minimal effect on any of the results in this paper.

5. Initial model calibration

I begin by fitting Models 1A and 1B to the data from the Beacon Heights core in order to obtain best estimates for muon interaction cross-sections. The reason to begin in this way is that this site has by far the lowest erosion rate, which, as discussed above, means that the parameter estimates will be least sensitive to both the accuracy of the erosion rate estimate from the surface nuclide concentration and the steadyerosion assumption. Borchers et al. (2016) estimated the erosion rate at Beacon Heights using the surface nuclide concentrations and spallogenic ¹⁰Be and ²⁶Al production rate estimates derived from independent calibration data, and obtained a range of erosion rate estimates between $7-15 \times 10^{-6}$ g cm⁻² yr⁻¹.

These erosion rates expressed in linear rather than mass units are 2.5-5.5 cm Myr⁻¹ 351 given rock density of 2.7 g cm⁻³. Note that throughout this paper I will use 2.7 g 352 cm⁻³ as a standardized value for rock density in converting mass erosion rates in g 353 cm⁻² yr⁻¹ to linear erosion rates in m Myr⁻¹, no matter what the actual density of 354 the lithology in question (for example, the actual density of the sandstone at Beacon 355 Heights is near 2.3 g cm⁻³). Because mass erosion rates in g cm⁻² yr⁻¹ are used 356 exclusively in all calculations in the paper, the value chosen for rock density has no 357 impact on the calculations, and the conversion to linear units is only important as a 358 means of representing erosion rates in more commonly used and more easily visualized 359 units. Choosing a standard density for the conversion avoids potential confusion about 360 whether surface erosion rates are computed for bedrock or soil densities, and simplifies 361 comparisons of erosion rates observed at different sites. 362

The large range of the erosion rate estimates from the spallogenic inventory at Bea-363 con Heights stems from differences in spallogenic production rate models. This is be-364 cause a corollary to the observation above that a production rate estimate is insensitive 365 to the assumed erosion rate when the erosion rate is low is that in the opposite situation, 366 when one seeks to estimate the erosion rate from the surface nuclide concentration, a 367 relatively small uncertainty in the production rate (in this case 10%) propagates into 368 a large uncertainty in the erosion rate (in this case 50%). However, Figure 1 shows 369 that a 50% uncertainty in the erosion rate, at an erosion rate of 3.5 cm Myr⁻¹, permits 370 better than 1% precision on estimates of muon production rates at all depths. Thus, the 371 steady-erosion assumption for this site contributes substantially less uncertainty to the 372 production rate estimates than the measurement uncertainties (> 3%). To summarize, 373 the erosion rate at Beacon Heights is low enough that uncertainty in the absolute mag-374 nitude and steadiness of the erosion rate does not contribute significant uncertainty to 375 estimates of muon production rates. In general, this is not the case for the other data 376 sets, so I conclude that the Beacon Heights data will yield the most accurate calibration 377 of muon interaction cross-sections or other fitted parameters. This, of course, was the 378 purpose of collecting the Beacon Heights core in the first place. 379

I now describe the fitting procedure for models 1A and 1B (those based on the Heisinger muon flux calculations). Here I (i) rewrite Equation 1 to separate production

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by neutron spallation, negative muon capture, and fast muon interactions; (ii) represent spallogenic production by a single exponential in mass depth; and (iii) follow Heisinger in representing production by each of the two muon production pathways as the product of a cross-section or likelihood with a stopping rate or a flux. This yields:

$$N_{i}(z) = \frac{P_{sp,i}e^{-z/\Lambda_{sp}}}{\lambda_{i} + \frac{\epsilon}{\Lambda_{sp}}} + f_{i}^{*} \int_{0}^{\infty} R_{\mu-}(z+\epsilon t)f_{C}f_{d}e^{-\lambda_{i}t}dt$$

$$+ \sigma_{0,i} \int_{0}^{\infty} \beta(z+\epsilon t)\Phi(z+\epsilon t)\bar{E}^{\alpha}(z+\epsilon t)N_{i}e^{-\lambda_{i}t}dt$$
(5)

In the first term, which describes spallogenic production, I take Λ_{sp} to be 140 g 386 cm⁻² for the Beacon Heights site (from Borchers et al., 2016) and ϵ to be 10×10^{-6} g 387 $cm^{-2} yr^{-1}$ (3.7 cm Myr⁻¹). As discussed above, for an erosion rate this low, estimates 388 of muon interaction cross-sections are insensitive to the assumed value of the erosion 389 rate; I duplicated the fitting exercise for erosion rates of 7 and 15×10^{-6} g cm⁻² yr⁻¹ 390 and verified that this range of erosion rates contributes less than a 1% variation in the 391 cross-section estimates. Given a prescribed value for ϵ , the integrals in the second 392 term (describing production by negative muon capture) and the third term (describing 393 production by fast muon interactions) are fully defined at any depth z by the Method 394 1A or 1B implementations of the Heisinger method (see the Heisinger papers for the 395 definition of the symbols). This leaves the surface production rate of nuclide *i* due to 396 spallation ($P_{sp,i}$; atoms g⁻¹ yr⁻¹), a negative muon capture probability for nuclide *i* (f_i^* ; 397 dimensionless), and a fast muon interaction cross-section for nuclide *i* ($\sigma_{0,i}$; barns) as 398 fitting parameters. Although we already have an estimate of $P_{sp,i}$ from Borchers, it is 399 not as precise as their estimate of Λ_{sp} , so I retain it as a nuisance parameter to limit the 400 effect of an inaccurate estimate of $P_{sp,i}$ on the muon interaction cross-sections. Also, 401 for computational practicality, I limit the integration time to eight half-lives of the 402 nuclide in question instead of infinity. To evaluate the integrals, I used the default nu-403 merical integration scheme in MATLAB R2015b. For Models 1A and 1B, an estimate 404 of the mean atmospheric pressure at the surface is needed. Here I use the Antarctic 405 pressure-elevation relationship of Stone (2000) to estimate this from the site elevation 406 (2183 m elevation implies 741.8 hPa). For Model 1B I assume zero geomagnetic cutoff 407 rigidity at the core site. To estimate best-fitting values of the free parameters, I mini-408 mized the chi-squared misfit between model and data, assuming reported measurement 409 uncertainties for the nuclide concentrations and disregarding uncertainties in the mass 410 depth of samples, and using the MATLAB implementation of the Nelder-Mead simplex 411 search method. 412

Figure 4 shows the results of fitting Equation 5 to the ¹⁰Be and ²⁶Al data from Bea-413 con Heights using the Model 1A code described above, that is, the MATLAB imple-414 mentation in Balco et al. (2008) of the Heisinger method. One aspect of the Heisinger 415 method is that the fast muon interaction cross-section σ is assumed to increase with 416 muon energy by a power law, such that if E is the muon energy, then $\sigma(E) = \sigma_0(E^{\alpha})$, 417 where σ_0 is the cross section for 1 GeV muons. Heisinger assumes α to be 0.75. Some 418 recent work (Lifton et al., 2014, J. Stone, written communications) has suggested that 419 the Heisinger method could be simplified without loss of accuracy by assuming $\alpha = 1$, 420



Figure 4: Fit of Model 1A with $\alpha = 0.75$ to ¹⁰Be (top) and ²⁶Al (bottom) measurements at Beacon Heights. Left panel shows all data on log-log axes; center panel shows data below 1000 g cm⁻² depth on semilog axes; right panel shows model-data residuals as ratio of measured to predicted concentration for each sample. Error bars in right panel are 1σ measurement uncertainty as reported in the source papers.

that is, a linear energy dependence for σ . Thus, I also show in Figure 5 the results of fitting Model 1A to the Beacon Heights data with $\alpha = 1$. Figure 6 shows the results of fitting Model 1B, which uses the implementation of Heisinger in Lifton et al. (2014) as described above, to the Beacon Heights data. Note that Model 1B also assumes $\alpha = 1$.



Figure 5: Fit of Model 1A with $\alpha = 1$ to ¹⁰Be (top) and ²⁶Al (bottom) measurements at Beacon Heights. Left panel shows all data on log-log axes; center panel shows data below 1000 g cm⁻² depth on semilog axes; right panel shows model-data residuals as ratio of measured to predicted concentration for each sample. Error bars in right panel are 1 σ measurement uncertainty as reported in the source papers.

Table 1 shows best-fitting muon interaction cross-sections for Models 1A (with $\alpha = 0.75$ and $\alpha = 1$) and 1B. All fit the data similarly, as expected given that they are based on the same physics. The scatter of the observations around the model predictions in all cases substantially exceeds reported measurement uncertainties. In other words, if we assume that the sole source of uncertainty is the reported measurement uncertainty in the ¹⁰Be and ²⁶Al concentrations, the probability-of-fit based on the chi-squared misfit is negligible, $< 1 \times 10^{-8}$ for either nuclide for any model. However, the



Figure 6: Fit of Model 1B to ¹⁰Be (top) and ²⁶Al (bottom) measurements at Beacon Heights. Left panel shows all data on log-log axes; center panel shows data below 1000 g cm⁻² depth on semilog axes; right panel shows model-data residuals as ratio of measured to predicted concentration for each sample. Error bars in right panel are 1σ measurement uncertainty as reported in the source papers.

distribution of residuals with respect to the model predictions is indistinguishable from 432 normal, unbiased, and displays no obvious trend (Figures 4-6), which suggests that the 433 main cause of scatter around the model predictions is unquantified measurement un-434 certainty rather than a systematic inaccuracy in the models. As the measurements from 435 this core are from three different laboratories, this agrees with the conclusions of Jull 436 et al. (2013) and Borchers et al. (2016) that measurement uncertainties for ¹⁰Be and 437 ²⁶Al reported by a particular laboratory underestimate true measurement uncertainties 438 in an interlaboratory-comparison sense. Borchers et al. (2016) performed a similar fit-439 ting exercise to the Beacon Heights data using code similar to that of Model 1B but 440 including expanded measurement uncertainties derived from the intercomparison of 441 Jull et al. (2013), and found acceptable probabilities-of-fit; I accept this result and have 442 not repeated this experiment here. Overall, these fitting exercises, as well as that of 443 Borchers et al. (2016), show that Models 1A and 1B both fit the Beacon Heights data 444 adequately. In addition, they indicate that simplifying the Heisinger method by setting 445 $\alpha = 1$ has little effect on agreement with observations; in fact, scatter of data around 446 model predictions is slightly (although not significantly) less for models with $\alpha = 1$ 447 than for Model 1A with $\alpha = 0.75$ (Table 1). Henceforth in this paper, therefore, I 448 assume $\alpha = 1$ when applying Model 1A. 449

The actual best-fitting values of f^* and σ_0 for Model 1A (with $\alpha = 1$) and Model 450 1B are similar to those obtained by Borchers et al. (2016), which, as noted above, used 451 code similar to Model 1B. As has been pointed out by numerous others previously 452 (e.g., Balco et al., 2008; Braucher et al., 2003, 2011, 2013), the cross-section estimates 453 for ¹⁰Be and ²⁶Al derived from fitting to available geological calibration data are $\sim 50\%$ 454 lower than the direct laboratory measurements of the cross-sections by irradiation of 455 synthetic mineral targets by high-energy muons reported by Heisinger (Table 1). In 456 other words, Heisinger's cross-section measurements overestimate observed ¹⁰Be and 457 ²⁶Al concentrations in this core by approximately a factor of two. Although it has 458 been proposed that this mismatch may be related to Heisinger's choice of a value for 459 α in scaling laboratory measurements of the fast muon interaction cross-section made 460 at high energy down to typical muon energies at moderate depths, this cannot fully 461 account for the mismatch. If, for example, one were to assume $\alpha \simeq 1.3$, this could 462 reconcile estimates of σ , but not f^* . 463

To summarize, muon production rates calculated using models based on Heisinger's 464 method with fitted muon interaction cross-sections successfully match ²⁶Al and ¹⁰Be 465 concentrations in the Beacon Heights deep core. This is a relatively stringent test be-466 cause the erosion rate is low enough that the comparison is insensitive to uncertainties 467 in the geological history of the site. However, calibrated values of f^* and σ_0 neces-468 sary to achieve this agreement are approximately 50% lower than values of the same 469 parameters measured by Heisinger in irradiation experiments. For the purposes of this 470 paper, this is interesting but not directly relevant: the aim here is not to figure out what 471 happened in Heisinger's experiments, but to accurately estimate production rates in 472 geologically useful settings. The fitting exercise shows that this can be accomplished 473 with the calibrated cross-section estimates. 474

I now turn to estimating the uncertainty in the production rate estimates. Note that for the purposes of this paper, the uncertainties in the cross-section estimates themselves are not of interest, because I am not interested in what the actual values of these

parameters are. Rather, I am interested in the uncertainties in estimates of production 478 rates based on best-fitting values of these parameters. Thus, I have not attempted to 479 compute an uncertainty in the cross-section estimates. Potentially, one can estimate 480 the uncertainty in model-predicted muon production rates by computing the scatter of 481 the residuals between predictions and observations for the calibration data. For data 482 from deeper than 1000 g cm⁻¹ where spallogenic production is negligible, this scat-483 ter for both Models 1A and 1B is $\sim 5\%$ for ¹⁰Be and $\sim 12\%$ for ²⁶Al (Table 1). This 484 only slightly exceeds the estimates of Borchers et al. (2016) for total measurement 485 uncertainty in ¹⁰Be and ²⁶Al of 3.6% and 10.1%, respectively, for low concentrations 486 representative of the deep samples at this site. Thus, as also noted by Borchers et al. 487 (2016) and already mentioned above, nearly all of the observed scatter around model 488 predictions is due to measurement uncertainty. Remaining differences between model 489 predictions and observations could, potentially, be due to (i) uncertainties in mass depth 490 estimates for the samples; (ii) possible changes in the erosion rate in the past, or (iii) 491 inaccuracies in the models. If we assume that of the observed 5% model-data scatter 492 for 10 Be, 3.6% is contributed by measurement uncertainty, this leaves a potential 3% 493 uncertainty due to these three factors. Thus, I conclude that uncertainty in muon pro-494 duction rates estimated by Models 1A and 1B at Beacon Heights that derives from the 495 models themselves is less than 3%. Note that this does not include any uncertainty 496 associated with scaling muon production rates from Beacon Heights to other locations. 497 I discuss that in the next section. 498



Figure 7: Fit of Model 2 to ¹⁰Be (left 2 panels) and ²⁶Al (right two panels) measurements at Beacon Heights. Left panel of each pair shows data below 1000 g cm⁻² depth on semilog axes; right panel of each pair shows model-data residuals as ratio of measured to predicted concentration for each sample. Error bars in right panel are 1σ measurement uncertainty as reported in the source papers.

Finally, I evaluate the exponential-approximation model for muon production rates 499 by fitting a single exponential function in depth, as proposed by various authors and 500 denoted 'Model 2' above, to the Beacon Heights data. Following the procedure given 501 in Braucher et al. (2013), I fit the equation $N_i(z) = N_{0,i} \exp(-\Lambda_{\mu,i}/z)$ to all data below 502 1000 g cm⁻² depth. This yields best-fitting values of a surface concentration of muon-503 produced nuclides $N_{0,i}$ (atoms g⁻¹ yr⁻¹) and an effective e-folding length $\Lambda_{\mu,i}$ (g cm⁻²). 504 The concentration $N_{0,i}$ can then be related to a surface production rate due to muons 505 $P_{0,i}$ using the relationship $P_{0,i} = N_{0,i}/(\lambda_i + \epsilon/\Lambda_{\mu,i})$ and an estimate of ϵ obtained from 506 the surface spallogenic nuclide inventory.

Figure 7 shows the result of fitting this model to the Beacon Heights data. It is 508 evident from this figure that this model is not adequate for fitting these data: a system-509 atic trend in the residuals, in which the model overestimates nuclide concentrations at 510 middle depths and underestimates them at the top and bottom of the core, is present 511 both for ¹⁰Be and ²⁶Al. In other words, a single exponential function fit to all data has 512 too long an e-folding length to fit the shallow data, and too short an e-folding length to 513 fit the deep data. This form of misfit is expected from the basic physics of muon pro-514 duction. Because the muon energy spectrum becomes more energetic with depth, the 515 instantaneous e-folding length of the production rate increases commensurately, and 516 this effect is not captured by a single exponential function. In addition, the e-folding 517 length $\Lambda_{\mu,i}$ obtained from fitting to the Beacon Heights data (2500 g cm⁻²; Table 1; 518 Figure 7) is substantially lower than values of 4000-5000 g cm⁻² obtained by Braucher 519 et al. (2013) from fitting to data from other sites. The primary reason for this difference, 520 and in addition the primary reason that the systematic misfit between the exponential 521 model and observations that is obvious at Beacon Heights is less evident in the data 522 considered by Braucher et al. (2013), is because that work considered data from sites 523 with erosion rates that are at one to two orders of magnitude higher than the erosion 524 rate at Beacon Heights. Heuristically, the effect of increasing the erosion rate is to in-525 crease the apparent e-folding length of the depth-concentration profile at a particular 526 depth by "dragging" the lower part of the profile up and increasing the importance of 527 the fast muon contribution (which has a longer e-folding length) at the expense of the 528 negative muon capture contribution (which has a shorter e-folding length). Thus, for 529 a site with a faster erosion rate, the depth-concentration profile conforms more closely 530 to that expected for fast muon production alone in that it displays an e-folding length 531 that is both longer and changes less rapidly with depth than would be the case for a 532 site with a slower erosion rate. The performance of an exponential approximation in 533 matching measured concentrations will therefore improve with increasing erosion rate. 534

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In Figure 8 I demonstrate this effect by computing expected concentrations of 535 muon-produced nuclides at moderate depths using Model 1A and then fitting them 536 with a single exponential function. In addition to the inverse relationship between ero-537 sion rate and fitted e-folding length for an exponential approximation noted above, fit-538 ted e-folding lengths are shorter at higher elevation, because production from stopped 539 negative muons (which has a shorter e-folding length) increases more rapidly with 540 elevation than production from fast muon interactions (which has a longer e-folding 541 length). This calculation correctly predicts fitted e-folding lengths in the 4000-5000 542 g cm⁻² range observed in the compilation by Braucher et al. (2013) at low elevations 543 and erosion rates of order 10 m Myr⁻¹, as well as the apparent e-folding length near 544 2500 g cm⁻² observed at Beacon Heights (Figure 7). Note, however, that it does not 545 explain the apparent e-folding length near \sim 5000 g cm⁻² for ¹⁰Be concentrations mea-546 sured at Cuiaba (Braucher et al., 2013); an apparent e-folding length closer to 3500 g 547 cm⁻² is expected for the relatively low erosion rate at that site. This mismatch may be 548 a consequence of the relatively high measurement scatter at that site, but it is otherwise 549 unexplained. It would be useful to revisit this issue by additional measurements at a 550 similar low-erosion-rate site at low elevation. 551

To summarize, the measurements at Beacon Heights, where the erosion rate is ex-



Figure 8: Calculation of apparent e-folding lengths for muon-produced ¹⁰Be concentrations as a function of erosion rate and elevation. This calculation uses Model 1A (with $\alpha = 1$) to predict steady-state nuclide concentrations between 1000-6000 g cm⁻² for a range of erosion rates, then fits an exponential function to the calculated nuclide concentrations in the same way as is shown for the Beacon Heights data in Figure 7 above. The left panel shows depth-concentration profiles for a range of erosion rates between 0.1-50 m Myr⁻¹ calculated with Model 1A as solid lines, with corresponding best-fitting exponential approximations as dashed lines. Higher erosion rates predict lower concentrations, so curves for higher erosion rates are at left in this figure. For higher erosion rates the contribution of negative muon capture production is less, which results in a longer apparent e-folding length and a closer match to an exponential approximation. At lower erosion rates, the contribution of negative muon capture proximation will be greater. The right-hand panel shows fitted e-folding lengths (that is, the slopes in semilog space of the red-lines in the left-hand figure) as a function of erosion rate. The red line shows results for 1013.25 hPa (sea level); blue dashed line is for 750 hPa (similar to the Beacon Heights site).

tremely low, highlight that an exponential approximation for muon production oversim-553 plifies the actual depth-dependence of muon production rates. A simplified exponential 554 model is not expected, based on theoretical considerations, to accurately predict muon 555 production rates or, therefore, to accurately predict concentrations of muon-produced 556 nuclides across a wide range in erosion rates. Observations at Beacon Heights agree 557 with this expectation. However, it is equally important to note that sites with an ero-558 sion rate this low are extremely rare on Earth, so it is highly likely, as pointed out 559 by Braucher et al. (2013) and many others, that the exponential model can potentially 560 provide an acceptably accurate and very much simpler method of estimating concen-561 trations of muon-produced nuclides for many geological applications. I explore that in 562 more detail later in the paper. 563

6. Geographic scaling

The previous section described fitting Models 1A and 1B to data from the Beacon Heights core and showed that both these models can be adequately fit to the observations. I now ask whether these models, which both include a geographic scaling method for the surface muon spectrum, are effective at scaling muon production rates from the calibration site at Beacon Heights to other locations.

I will not further discuss Model 2 (the single-exponential model) in this section, because, as discussed above, it is evident from a comparison of Figures 7 and 8 above

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with the results of Braucher et al. (2013) that Model 2 with parameters calibrated from 572 Beacon Heights will perform poorly at other bedrock core sites with different ero-573 sion rates. The fitted e-folding length for subsurface nuclide concentrations at Beacon 574 Heights (2500 g cm⁻²; see Table 1) is very different from that observed at La Ciotat 575 and Leymon Quarry (4000-5000 g cm⁻²; see Table 2 of Braucher et al., 2013), so an 576 exponential model derived from the former would clearly not fit the latter. Thus, I have 577 not carried out this exercise. I return to the question of when a simplified exponential 578 model can be used with acceptable accuracy for geological applications later in the 579 paper.

6.1. Bedrock core data at relatively high-erosion-rate sites

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Ideally, one could test geographic scaling models by calibrating the models at one 582 location with an extremely low erosion rate (e.g., Beacon Heights), using the calibrated 583 models to predict nuclide concentrations at a different site with an equally low erosion 584 rate, and comparing these predictions to measurements at the new site. If the erosion 585 rate at the new site were low enough, we would not need an accurate estimate of the 586 erosion rate. However, other sites where data from deep bedrock cores are available 587 (La Ciotat and Leymon Quarry) have surface erosion rates (inferred from spallogenic 588 nuclide concentrations in surface samples) that are two orders of magnitude higher, in 589 the 10-40 m Myr⁻¹ range. At these erosion rates, as discussed above and shown in 590 Figure 1, the concentration of muon-produced nuclides has a strong dependence on 591 the erosion rate. At these higher-erosion-rate sites, we can estimate the steady-state 592 erosion rate from the spallogenic nuclide concentration and use that erosion rate with 593 a muon production model to predict concentrations of muon-produced nuclides in the 594 subsurface for comparison with observations. However, the results of this compari-595 son are very dependent on the steady-erosion assumption, because if the erosion rate 596 changed over time, the spallogenic nuclide inventory would reach equilibrium with a 597 new erosion rate faster than the muon-produced inventory. In this case, there would 598 be no single value for a steady-state erosion rate that could successfully predict both 599 spallogenic nuclide concentrations in near-surface samples and muon-induced nuclide 600 concentrations in deeper samples. We could, therefore, observe a mismatch between 601 predicted and measured concentrations of muon-produced nuclides either because (i) 602 the muon production rate scaling model is incorrect, or (ii) the scaling model is cor-603 rect, but our assumption that both spallogenic and muon-produced nuclide inventories 604 are in equilibrium with the same steady erosion rate is incorrect. In practice, there is 605 no way to independently verify the steady-state assumption, which then implies that 606 we can not evaluate the scaling model with high confidence either. In fact, one could 607 argue that the steady-state assumption is likely to be false because all the sites have 608 experienced periodic climate change, which presumably affected erosion rates at least 609 to some extent, during the Pleistocene (see Braucher et al., 2003, for a more detailed 610 discussion of this effect). 611

Despite the difficulty of differentiating the performance of muon production models from the validity of the steady state assumption at these sites, this test is still potentially useful, because one might find that the muon scaling models calibrated at Beacon Heights were not able to successfully predict observed concentrations of muonproduced nuclides at the other core sites for any erosion rate. If the production model could not successfully predict observed nuclide concentrations for any value of the erosion rate, this would show that the model is incorrect. In other words, although it is not possible to prove unambiguously that a scaling model is correct, it might be possible to show that it is incorrect. Thus, in this section I carry out this comparison exercise for the La Ciotat, Leymon Low, and Leymon High cores. I address the data from excavations at Cuiaba and Macraes Flat in the next section. Here I am fitting the following equation to the ²⁶Al and ¹⁰Be data (separately) at each site:

$$N_{i}(z) = N_{sp}e^{-\frac{z}{\Lambda_{sp}}} + f_{i}^{*} \int_{0}^{\infty} R_{\mu-}(z+\epsilon t)f_{C}f_{d}e^{-\lambda_{i}t}dt$$

$$+ \sigma_{0,i} \int_{0}^{\infty} \beta(z+\epsilon t)\Phi(z+\epsilon t)\bar{E}^{\alpha}(z+\epsilon t)N_{i}e^{-\lambda_{i}t}dt$$
(6)

This is the same as Equation 5 except that the spallogenic nuclide inventory, which 624 is defined as a function of the erosion rate in Equation 5, is represented here by a sin-625 gle parameter N_{sp} with units of atoms g⁻¹. The muon interaction cross-sections f_i^* 626 and $\sigma_{0,i}$ are known from fitting each model to the Beacon Heights data as described 627 above, and the parameter Λ_{sp} is obtained for the elevation and magnetic cutoff rigidity 628 of each site using the table of Marrero et al. (2016) (magnetic cutoff rigidity is esti-629 mated from the geographic latitude of the site using the equation in Lifton et al., 2014). 630 This leaves the erosion rate ϵ and the spallogenic nuclide inventory at the surface N_{sp} 631 as the fitting parameters. Subsequently, given an (independent) estimate of the nuclide 632 production rate due to spallation (P_{sp}) at the site, one can then interpret N_{sp} as an ap-633 parent steady-state erosion rate implied by the spallogenic nuclide inventory ϵ_{sp} using 634 the relationship $N_{sp} = P_{sp}/(\lambda + \epsilon_{sp}/\Lambda_{sp})$. Basically, the purpose of this procedure is 635 to make two estimates of the steady-state erosion rate at each site independently from 636 the spallogenic and muon-produced nuclide inventories. I then consider (i) whether 637 the muon scaling model can fit the measurements for any value of the erosion rate, 638 and (ii) whether estimates of the erosion rate from spallogenic and muon-produced nu-639 clide concentrations agree. As discussed above, agreement between the two erosion 640 rate estimates would, in a general sense, indicate that the muon production model is 641 successful at scaling muon production rates between sites, but disagreement between 642 the estimates could mean either that the production model fails or that the steady-state 643 erosion assumption fails. In other words, a failure to fit the data at any erosion rate 644 could falsify the hypothesis that a muon scaling model is correct, but success in fitting 645 the data would not prove the hypothesis. Other details of the fitting procedure here are 646 as follows. Following the discussion above, in this section I will only consider the case 647 where $\alpha = 1$ for both Model 1A and Model 1B. I estimate the atmospheric pressure 648 at each site using the ERA40 atmosphere as implemented by Lifton et al. (2014). To 649 compute spallogenic production rates necessary for interpreting the best-fitting spallo-650 genic inventory N_{sp} as an erosion rate, I use the 'St' scaling method and production 651 rate calibration of Borchers et al. (2016). 652

Figures 9 and 10 show the results of this fitting exercise. Scatter around the bestfitting models at most sites is greater than observed at Beacon Heights (Table 2), but no systematic bias to the residuals is evident. Model 1A displays slightly less scatter



Figure 9: Best fit of Models 1A (with $\alpha = 1$) and 1B, with cross-sections calibrated to the Beacon Heights data, to ¹⁰Be measurements for La Ciotat and Leymon Quarry bedrock cores. Left panel of each pair shows data below 1000 g cm⁻² depth on semilog axes with best-fitting predicted nuclide concentrations for Model 1A (solid line) and 1B (dashed line). Right panel of each pair shows model-data residuals as ratio of measured to predicted concentration for each sample: solid symbols are for Model 1A and open symbols for Model 1B. Error bars in right panel are 1σ measurement uncertainty as reported in the source publications. Although only the data below 1000 g cm⁻² are shown here, the models were fit to all data.



Figure 10: Best fit of Models 1A (with $\alpha = 1$) and 1B, with cross-sections calibrated to the Beacon Heights data, to ²⁶Al measurements for La Ciotat and Leymon Quarry bedrock cores. Axes and symbols are the same as for Figure 9.

with respect to the observations than Model 1B in all cases, but the differences are not significant.

Table 2 shows apparent steady-state erosion rates inferred from spallogenic and muon-produced nuclide inventories for ²⁶Al and ¹⁰Be for these three sites. At La Ciotat, erosion rate estimates from the muon-produced inventory for ²⁶Al and ¹⁰Be and for both models (27-33 m Myr⁻¹) are substantially (\sim 50%) higher than the erosion rate estimated from the spallogenic inventory (18-21 m Myr⁻¹). As noted above, this difference could be explained either by a failure of the muon production scaling or by the fact that the site has not experienced steady erosion; it is not possible to determine which from these data alone.

At the Leymon High site, on the other hand, erosion rate estimates from the muonproduced inventory (12-14 m Myr⁻¹) are lower, although not significantly so, than erosion rate estimates from the spallogenic inventory (17 m Myr⁻¹). At the Leymon Low site, the two estimates are similar $(8-13 \text{ m Myr}^{-1})$. The agreement between erosion rate estimates from spallogenic and muon-produced inventories at this site is consistent with the hypothesis that (i) nuclide concentrations have reached steady state with the erosion rate and also (ii) the muon scaling models are correct. However, again, it is not possible to exclude the possibility that both parts of the hypothesis are incorrect in such a way as to produce spurious agreement between the erosion rate estimates.

To summarize, it is possible to find erosion rates that provide a good fit to core 675 data at La Ciotat and Leymon Quarry using either Model 1A or 1B calibrated with the 676 Beacon Heights data. Thus, these data provide no evidence that either of these mod-677 els is incorrect or inaccurate. However, because of the ambiguity regarding whether 678 poor agreement between steady erosion models inferred from spallogenic and muon-679 produced nuclides indicates a failure of the scaling model or a failure of the steady-state 680 assumption, it is not possible to evaluate the performance of the scaling models any 681 more precisely than this with these data. In general, however, results from these three 682 core sites show better agreement between apparent erosion rates derived from muon-683 produced and spallogenic nuclide inventories at lower-erosion-rate sites. This would tend to indicate that mismatches between erosion rate estimates are more likely caused 685 by non-steady-state erosion than by errors in muon production rate scaling. To summa-686 rize, data from these bedrock cores are consistent with, although they do not prove, the 687 hypothesis that both Model 1A and 1B are accurate for scaling muon production rates 688 from Beacon Heights to these sites.

6.2. Other sites with lower erosion rates

Two other sites not discussed in the previous section, Macraes Flat and Cuiaba, are 691 potentially more useful for testing the geographic scaling in Models 1A and 1B because these sites are deeper (Macraes Flat) or have a lower erosion rate (Cuiaba) than 693 the La Ciotat or Leymon Quarry sites. Both of those factors reduce the dependence of the muon-produced nuclide concentrations on the erosion rate. Thus, we can use some data from these sites to test the muon scaling models with less reliance on the assumption that spallogenic and muon-produced erosion rates are in equilibrium with the same steady erosion rate. The Cuiaba site is potentially particularly useful for this, because its low-elevation, low-latitude location maximizes the difference in production rate scaling relative to the Beacon Heights site. On the other hand, both Macraes Flat

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and Cuiaba are geomorphically more complicated than the bedrock core sites. In both cases, samples were collected opportunistically from existing mining excavations, and a number of assumptions about the mass thickness of material removed during various stages of mine site preparation, etc., were required to estimate the sample depths below the surface. This contributes some uncertainty to using these measurements as calibration data.

6.2.1. Cuiaba

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At Cuiaba, Braucher et al. (2003) collected samples from two sites: a deep open-708 pit excavation where the original soil surface was not present and the amount of pre-709 stripping prior to excavation was not exactly known; and a second site, in a different 710 location from the main excavation, where the original soil surface was present but only 711 a shallow excavation could be made. This situation creates two potential uncertainties: 712 first, in reconstructing the thickness of soil originally present at the deep site; second, 713 in the assumption that the erosion rate is the same at both sites. These authors inferred 714 the original depth of the deep samples by comparing measured ¹⁰Be concentrations at 715 both sites and choosing an original soil thickness at the deep site that was consistent 716 with nearby observations from undisturbed soils and also yielded a continuous nuclide 717 concentration profile when data from both sites were considered together. I use a sim-718 ilar strategy here. First, I fit Equation 6 to the ¹⁰Be measurements from both sites, 719 with the original cover thickness above the deep site as an additional free parameter. 720 This yielded a best-fitting cover thickness of 1490 g cm⁻² (for Model 1A with $\alpha = 1$), 721 which is similar to the results of Braucher et al. (2003) (although this value is only 722 stated as "about 4 m" in the paper, examination of their Figure 2 indicates it is near 723 1200 g cm⁻²), and a steady erosion rate near 0.5 m Myr⁻¹. Second, using this value of 724 the cover thickness, I compare the measurements to predicted ¹⁰Be concentrations for 725 a range of erosion rates spanning an order of magnitude between 0.2-2 m Myr⁻¹ using 726 Models 1A and 1B. 727

Figure 11 shows this comparison. Although the measurements at this site are some-728 what scattered, if we accept that nuclide concentrations at the site have, in fact, reached 729 equilibrium with a steady erosion rate somewhere in the range 0.2-2 m Myr⁻¹, ¹⁰Be 730 concentrations predicted by Model 1A (with $\alpha = 1$) for this range of erosion rates are 731 consistent with the observations in that there is no significant systematic residual be-732 tween observations and predictions. However, concentrations predicted by Model 1B 733 systematically underestimate the observed concentrations for any erosion rate. This 734 would imply that the geographic scaling of muon fluxes in Model 1B overestimates 735 the difference in subsurface muon flux between Beacon Heights and Cuiaba. Note that 736 the mismatch between Model 1B and the observations cannot be explained by a po-737 tential inconsistency in using Model 1A to estimate the original soil thickness at the 738 deep site: the soil thickness would have to be less than 500 g cm⁻² to explain the mis-739 match, which appears inconsistent with the authors' observations as described in the 740 paper. A failure of the steady-erosion assumption could not explain it either, because 741 the observed concentrations are systematically higher than Model 1B would predict 742 even for zero erosion at this site. To summarize, although the difficulty in estimating 743 the original configuration of the site creates some uncertainties in this conclusion, this 744 comparison between model predictions and the obervations at Cuiaba indicates that the 745



Figure 11: Comparison of ¹⁰Be concentrations at Cuiaba site with those predicted by Models 1A (with $\alpha = 1$; solid lines) and 1B (dashed lines), with cross-sections calibrated to the Beacon Heights data as described above, for a range of erosion rates. Lines correspond to erosion rates of 0.2 (rightmost), 0.5, 1, and 2 (leftmost) m Myr⁻¹. Error bars show 1σ uncertainties as reported in the source paper.

simpler elevation scaling of the muon flux in Model 1A is more accurate than the more 746 complicated geographic scaling of Model 1B in scaling production rates calibrated at 747 Beacon Heights to this low-elevation, low-latitude site. 748

6.2.2. Macraes Flat

At Macraes Flat, the likely erosion rate is higher, but samples were collected from 750 deeper below the surface. Fitting Equation 6 to subsurface data with either Model 1A 751 or 1B yielded erosion rate estimates of 5-10 m Myr⁻¹, in approximate agreement with 752 the estimate of 12 m Myr⁻¹ by Kim and Englert (2004). This site is composed of sev-753 eral distinct excavations, each of which represents a different, non-overlapping, range 754 of depths below the original surface. In addition, these authors concluded from both 755 geomorphic criteria and the observations themselves that the near-surface concentra-756 tions are not in equilibrium with steady erosion. However, nuclide concentrations in 757 samples from the deepest site (> 10,000 g cm⁻²) are relatively insensitive both to the 758 steady-erosion assumption and to potential inaccuracies in estimating the original land 759 surface configuration. Thus, I carried out an analysis similar to that described above 760 for the Cuiaba site for these deepest samples.

Figure 12 shows the results. Predictions for a range of erosion rates between 2 and 762 20 m Myr⁻¹ are consistent with ¹⁰Be and ²⁶Al measurements between 10,000-25,000 g 763 cm^{-2} depth. However, the deepest ¹⁰Be measurement at 50,000 g cm⁻² depth is almost 764 double the concentration predicted at that depth by either scaling model. The signifi-765 cance of the failure to fit with the deepest ¹⁰Be measurement is unclear; the only way 766 to reconcile this measurement with the ¹⁰Be concentrations in the overlying samples 767 would be if α was substantially greater than 1, which would be inconsistent with other 768 calibration data as well as with laboratory measurements. It would probably be valu-769 able to replicate this ¹⁰Be measurement. The corresponding ²⁶Al measurement is not 770

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Figure 12: Comparison of ¹⁰Be and ²⁶Al concentrations at Macraes Flat with those predicted by Models 1A (with $\alpha = 1$; solid lines) and 1B (dashed lines), with cross-sections calibrated to the Beacon Heights data as described above, for a range of erosion rates. Error bars show 1σ uncertainties as reported in the source paper. Lines correspond to erosion rates of 2 (rightmost), 5, 10, and 20 (leftmost) m Myr⁻¹.

similarly anomalous, although it has relatively large measurement uncertainty. Model 771 1A and 1B predictions are very similar at this site; both can explain the measurements in the 10,000-25,000 g cm⁻² depth range given an erosion rate in the range 2-20 m 773 Myr^{-1} .

6.3. Summary: geographic scaling of muon production rates

In the previous section I calibrated muon production models using data from the 776 lowest-erosion-rate site available at Beacon Heights. In this section, I carried out a 777 series of tests to determine whether the calibrated models can successfully account for 778 ¹⁰Be and ²⁶Al concentrations in subsurface calibration samples elsewhere. The only 779 significant difference in the performance of model 1A and 1B is that Model 1B appears 780 to overestimate the difference in production rate scaling between the high-elevation, 781 high-latitude (e.g., low geomagnetic cutoff rigidity) at Beacon Heights and the low-782 elevation, low-latitude (high cutoff rigidity) site at Cuiaba. At all other sites, both 783 models 1A and 1B are successful in matching observations.

7. Muon production of ¹⁴C in quartz

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For completeness and to facilitate use of Models 1A and 1B to compute ¹⁴C pro-786 duction rates due to muons, I also fit Models 1A and 1B to measurements of in-situ-787 produced ¹⁴C in quartz in the Leymon High core described by Lupker et al. (2015), 788 using a similar approach as described above for fitting production models to ²⁶Al and 789 ¹⁰Be measurements in the Beacon Heights core. Although the erosion rate is much 790 higher at the Leymon High site (10-15 m Myr⁻¹; see Table 2), the short half-life of 791 ¹⁴C improves the precision of production rate estimates for a given erosion rate by two 792 orders of magnitude relative to ¹⁰Be or ²⁶Al. The analysis shown in Figure 1, if applied 793 to ¹⁴C measurements, shows that the expected uncertainty in calibrating muon produc-794 tion models for ¹⁴C at the Leymon High core is similar to that in calibrating production 795 models for ²⁶Al and ¹⁰Be at Beacon Heights.

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The limited depth range of the ¹⁴C measurements at the Leymon High core, as 797 noted by Lupker et al. (2015), means that these data provide very little resolving power 798 on the fast muon interaction cross-section σ_0 , so it is only possible to accurately es-799 timate the negative muon capture probability f_{14}^* . Thus, I fit Equation 5 to the ¹⁴C 800 measurements by assuming that the erosion rate is 12 m Myr⁻¹ as inferred from ¹⁰Be 801 and ²⁶Al concentrations, assuming $\Lambda_{sp} = 160$ g cm² (Lupker et al., 2015), and assum-802 ing the experimentally determined value of σ_0 from Heisinger et al. (2002b, ; 2.4 μ b 803 for $\alpha = 1$). This leaves two free parameters: f_{14}^* and a production rate due to spallation 804 P_{sp} , which is a nuisance parameter for the present purpose. 805



Figure 13: Fit of Model 1A with $\alpha = 1$ (top) and Model 1B (bottom) to ¹⁴C measurements at Leymon High core. Left panels show all data on log-log axes; center panels show data below 1000 g cm⁻² depth on semilog axes; right panels show model-data residuals as ratio of measured to predicted concentration for each sample. Error bars are 1σ 'total' uncertainties of Lupker et al. (2015).

Figure 13 shows the results of this fitting exercise. This yields best-fitting values 806 for f_{14}^* of 0.116 (Model 1A) and 0.114 (Model 1B). This closely replicates the results 807 of a similar fitting exercise by Lupker et al. (2015), although they used slightly different 808 code to compute the muon fluxes and in addition allowed σ_0 to float in their calculation, 809 and obtained a slightly different best-fitting value as a result (0.134). As also noted by 810 these authors, all these estimates are in much better agreement with the experimentally 811 determined value of Heisinger (0.137) than is the case for 10 Be or 26 Al. 812

8. Needed accuracy/precision for burial-dating applications.

The previous sections show that, as expected from theoretical considerations, rela-814 tively complicated methods for estimating subsurface production rates that are based on 815 downward propagation of the surface muon flux perform better at matching available 816 calibration data than simplified exponential approximations. In this and subsequent 817 sections, I discuss to what extent this difference is important for various geological 818 applications that require computing production rates due to muons.

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First, I consider the related applications of burial dating and depth profile dating. 820 Burial dating relies on measurements of pairs of cosmic-ray-produced nuclides that 821 are produced in a fixed ratio during surface exposure, but have different half-lives (see 822 Granger, 2006, for a review). In a sample that is exposed at the surface, the ratio of 823 these nuclides will conform to the production ratio. If the sample is then buried so 824 that it is shielded from the cosmic-ray flux, the two nuclides will decay at different 825 rates, so their ratio will be related to the duration of burial. Although this basic concept 826 does not inherently require one to compute subsurface production rates due to muons, 827 in many practical applications of this method samples are buried at relatively shallow 828 depths, so nuclide production continues after burial, although at a lower rate than at the 829 surface. In this case, determining the burial age of the sample requires an estimate of 830 the production rate due to muons at the depth the samples are buried, and the accuracy 831 of this estimate of the subsurface production rate has a significant effect on the accuracy 832 of the eventual burial age. How important this effect is depends on the relationship 833 of the total nuclide concentrations to the subsurface production rates. It is relatively 834 unimportant if the sample was buried deeply with a high nuclide concentration and/or 835 has been buried for a short time; it is very important if the sample has been buried 836 for a long time at a shallow depth, so that the measured nuclide concentrations are 837 predominantly the result of post-burial production. 838

Depth profile dating is an approach aimed at determining both the exposure age 839 and erosion rate of landforms that have eroded after their initial emplacement (e.g., 840 see summary in Hidy et al., 2010). It relies on the fact that the spallogenic nuclide 841 inventory near the surface of an eroding landform is strongly dependent on the erosion 842 rate, and weakly dependent on the exposure age, whereas the muon-produced nuclide 843 inventory at several meters depth is weakly dependent on the surface erosion rate and 844 strongly dependent on the exposure age. Thus, measuring the nuclide concentration in 845 both surface and subsurface samples, in theory, allows one to estimate both the expo-846 sure age and the erosion rate. This application, again, requires accurate estimates of 847 subsurface production rates due to muons, and, in general, the accuracy of the expo-848 sure age estimate scales directly with the accuracy of the subsurface production rate 849 estimate. That is, a 20% uncertainty in the subsurface production rate estimate will directly scale into a 20% uncertainty in the derived exposure age.

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The common feature of burial-dating and depth-profile dating applications is that, in most cases, samples collected in the subsurface have been at close to the same depth since they were emplaced. Thus, these applications require accurate estimates of the production rate due to muons at specific depths, which, in turn, requires one of the methods based on downward propagation of the surface muon spectrum, that is, Method 1A or Method 1B. As expected from physical principles and as shown by the fact that a simple exponential approximation for the production rate due to muons cannot be fit to calibration data from sites that span a range of erosion rates, a single global exponential approximation is not adequate for predicting production rates at arbitrary location and depth.

However, it is also important to note that in most burial-dating and depth-profile 862 dating applications, it is only necessary to compute subsurface production rates for a 863 relatively small depth range at a particular site. For example, if one collects samples 864 between 500 and 1000 g cm⁻² depth in a landform that is 100,000 years old and erod-865 ing at no more than 4×10^{-4} g cm⁻² yr⁻¹ (that is, 2 m Myr⁻¹ for alluvial sediment with 866 density 2 g cm⁻³), it is only necessary to estimate production rates in the range 500-867 1040 g cm⁻². Given a restricted depth range as in this example, one can construct a 868 site-specific exponential approximation to subsurface production rates computed using 869 Model 1A or 1B, and use that approximation to simplify further calculations without 870 significant loss of accuracy. Figure 14 shows an example that uses a sum of two expo-871 nential functions; other examples include Granger and Smith (2000) and Balco et al. 872 (2011a), who used sums of three exponential functions to approximate muon produc-873 tion rates at specific sites. A similar approach is described by Stone et al. (1998), who 874 used polynomial functions in $\log(z)$ for this purpose. To summarize, the important 875 point here is that in burial-dating and depth-profile applications when one only needs 876 information about muon production rates in a relatively small depth range, one can use 877 a site-specific exponential approximation withouth loss of accuracy. However, an ex-878 ponential approximation accurate at one site will not be accurate for any other location, 879 except by accident. 880

Finally, I attempt to estimate the total uncertainty in estimating the nuclide pro-881 duction rate due to muons at a arbitrary location and depth using Model 1A or 1B (or 882 a site-specific approximation thereto). Obviously, this estimate is necessary for com-883 puting the contribution of uncertainty in muon production rate estimates to the total 884 uncertainty in a burial or depth-profile age. How to estimate this uncertainty, however, 885 is not at all obvious. Simply attempting to compute an overall misfit or scatter between 886 measurements and predictions for all the calibration sites would likely be incorrect 887 because of the ambiguity in whether agreement or disagreement between model calcu-888 lations reflects accuracy or inaccuracy in the steady-erosion assumption, or the production model. In many cases one could compensate for a failure of the scaling model by 890 adjusting the erosion rate to obtain a good fit between predictions and measurements 891 regardless.

One can potentially obtain an upper limit on the global uncertainty in predicting 893 muon production rates for arbitrary depth and location by comparing model predic-894 tions calibrated at Beacon Heights to observations at Cuiaba, which is the site besides 895



Figure 14: Example of a site-specific exponential approximation to muon production rates computed using Model 1A over a limited depth range. Circles are ¹⁰Be production rates computed using Model 1A for depths between 400 and 1200 g cm⁻² at 1000 m elevation; black line is a best-fitting sum of two exponentials (specifically, $P_{mu} = 0.0413 \exp(-z/3264) + 0.0669 \exp(-z/811)$). Scatter of the Model 1A production rate estimates around the approximation is less than 0.1%.

Beacon Heights at which the steady-erosion assumption is least important (and in ad-896 dition it is the site that has the largest scaling difference from Beacon Heights). The 897 data from Cuiaba show scatter well in excess of measurement uncertainty in relation 898 to any model prediction that is smooth in depth - the scatter around any exponential or 899 linear fit to the data from the deep site is at least 30%, compared to mean measurement 900 uncertainty near 15% - so it is hard to evaluate any model fit purely on the basis of this 901 scatter metric. A potentially more useful strategy would be to try to put an upper bound 902 on the total scaling uncertainty for the muon production rate models by assuming that 903 the erosion rate at Cuiaba is 0.8 m Myr⁻¹, as inferred from the spallogenic nuclide 904 inventory at the surface. If we use Model 1A to compute subsurface nuclide concentra-905 tions on this basis, we find that predicted ¹⁰Be concentrations underestimate measured 906 concentrations at depths >1000 g cm⁻² by 25% (\pm 28%). Of course, a 25% difference 907 is not significant at high confidence given 28% scatter, and in addition this estimate de-908 pends on the amount of soil cover assumed for the lower site at Cuiaba. If we reduce the 909 assumed soil thickness to 900 g cm⁻², which is likely permitted based on the published 910 description of the site, there would be zero systematic offset between the measurements 911 and the Model 1A predictions (Model 1B would still underestimate measurements by 912 30%), but on the other hand this would reduce agreement between ¹⁰Be concentrations 913 at the deep and shallow sites. Overall, however, I argue from this reasoning that 25% 914 is an upper bound for the total uncertainty in predicting ¹⁰Be and ²⁶Al production rates 915 due to muons at arbitrary location and depth using Model 1A (the corresponding figure 916 for Model 1B is 50%). In other words, the 25% difference between predictions and 917 observations at this site as calculated above is due to (i) inaccuracy in estimating the 918 soil cover thickness; (ii) possible unsteady erosion; (iii) measurement uncertainty; and 919 (iv) inaccuracy in the muon production rate scaling. Thus, an upper bound on the un-920

certainty in (iv), the muon production rate scaling, must be less than 25% (for Model 1A) or 50% (for Model 1B). It is, in addition, certainly possible – and it is impossible to disprove – that total scaling uncertainty for muon production using Model 1A is no 923 larger than scaling uncertainty for spallogenic production, which has been estimated from scatter in surface production rate calibration data to be in the range of 6-10% (Borchers et al., 2016).

To summarize, I propose that the total uncertainty in computing subsurface muon production rates at arbitrary location and depth using Model 1A is less than 25% un-928 der any circumstances and, by analogy with better estimates for scaling accuracy for 929 spallogenic production, likely 10% or better in most cases (i.e., where the scaling dif-930 ference between Beacon Heights and the site of interest is much smaller than that be-931 tween Beacon Heighs and Cuiaba). On the other hand, available evidence indicates the 932 uncertainty in computing production rates using Model 1B may be larger, possibly as 933 much as 50%. Calibration data are not sufficient to make a more precise estimate of 934 the global scaling uncertainty for these production models; the most efficient way to 935 address this problem would most likely be to collect additional calibration data from a 936 bedrock core at a low-elevation, low-erosion-rate site similar to Cuiaba.

9. Needed accuracy/precision for surface exposure dating applications.

Surface exposure dating is a much less demanding application of muon production 939 rate calculations for two reasons. First, muon production is a small fraction of total 940 surface production. Second, samples that can be accurately exposure-dated are by def-941 inition at sites where the surface erosion rate is very low, so it is not necessary to have 942 any information about the production rate below a few centimeters depth. Assume, 943 for example, that surface production due to muons is 2% of total surface production. 944 If we accept the argument above that muon production rates calculated using Model 945 1A calibrated at Beacon Heights are accurate to better than 25% in all cases, this im-946 plies that the uncertainty in estimating the total surface production rate at an arbitrary 947 location that is attributable to scaling uncertainty in Model 1A is only 0.5%. This is 948 substantially less than the uncertainty in scaling spallogenic surface production, which 949 is generally believed to be at least ~6% (Borchers et al., 2016). Note that this is not 950 necessarily true for exposure-dating using in-situ-produced ¹⁴C in quartz: production 951 of ¹⁴C by muons accounts for up to 20% of total surface production. Thus, for ¹⁴C, a 952 25% uncertainty in estimating muon production rates would equate to a 5% uncertainty 953 in the total surface production rate estimate. On the other hand, this indicates that sur-954 face production rate calibration data for ¹⁴C can potentially provide some constraints 955 on the uncertainty in scaling muon production rates. However, I have not pursued that 956 approach further here. 957

To summarize, Model 1A appears to have substantially better than needed accuracy for exposure-dating applications using ¹⁰Be and ²⁶Al. Of course, Model 1A is also computationally quite complicated. If it is also more precise than necessary, this suggests that a much simpler approximation of surface production rates due to muons would be adequate for ¹⁰Be and ²⁶Al exposure dating. Surface production rates predicted by Model 1A can be accurately approximated by:

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$$P_{\mu,i,0}(h) = P_{\mu,SL,i} \exp\left[\frac{(1013.25 - h)}{L_i}\right]$$
(7)

Here $P_{\mu,i,0}(h)$ is the total (including both negative muon capture and fast muon in-964 teractions) surface production rate of nuclide i due to muons (atoms $g^{-1} yr^{-1}$) at atmo-965 spheric pressure h (hPa), $P_{\mu,SL,i}$ is a nominal surface production rate (atoms g⁻¹ yr⁻¹) 966 of nuclide *i* due to muons at sea level (1013.25 hPa), and L_i is an effective e-folding 967 length in atmospheric pressure for nuclide i (hPa). Note that although atmospheric 968 attenuation of the muon flux is independent of the nuclide being produced, it is de-969 pendent on muon energy. Because the proportion of production due to negative muon 970 capture and fast muon interactions varies by nuclide, production of each nuclide in-971 volves a different part of the muon energy spectrum, and the value of L_i in this formula 972 is specific to each nuclide. As shown in Figure 15, Equation 7 can be fit to surface 973 production rates predicted by Model 1A quite accurately; the maximum difference be-974 tween the Model 1A predictions and the approximation with best-fitting parameters in 975 the range 450-1013.25 hPa is 2% (¹⁰Be) and 1% (²⁶Al). The best fitting parameters are $P_{\mu,SL,10} = 0.0735$ atoms g⁻¹ yr⁻¹, $P_{\mu,SL,26} = 0.6764$ atoms g⁻¹ yr⁻¹ $P_{\mu,SL,14} = 3.067$ atoms g⁻¹ yr⁻¹, $L_{10} = 299.2$ hPa, $L_{26} = 288.0$ hPa, and $L_{14} = 267.8$ hPa. This is 976 977 978 very similar to the elevation scaling formulae proposed by Braucher et al. (2013) and 979 Lal (1991), and highlights the fact that although simple exponential approximations for 980 muon production rate scaling are not globally accurate for arbitrary location and depth, 981 they are more than accurate enough for some geological applications. 982



Figure 15: Circles are total surface production rates due to muons predicted by Model 1A (with $\alpha = 1$) as calibrated at Beacon Heights for ¹⁰Be (top), ²⁶Al (middle), and ¹⁴C (bottom). Lines show Equation 7 fit to the Model 1A predictions.

To summarize, Model 1A has better accuracy than needed in estimating surface 983

production rates due to muons for exposure-dating applications using ²⁶Al and ¹⁰Be, and surface production rates computed by Model 1A can be fit very precisely by a single exponential function in atmospheric pressure. Thus, the simple approximation in Equation 7 is accurate enough for exposure-dating applications. If this had been recognized by Balco et al. (2008), they could have saved a lot of computation time in exposure-dating calculations.

Note that for exposure-dating applications it is also necessary to have some infor-990 mation about the depth-dependence of production due to muons, because computing 991 the exposure age of sites with non-negligible surface erosion rates requires integrat-992 ing production rates with respect to depth. For typical exposure-dating applications 993 at sites with exposure ages of tens of thousands of years, surface erosion rates are in 994 the range of millimeters per thousand years or lower. Thus, for most exposure-dating 995 applications, samples have not been more than ~10 cm below the surface during the 996 period of exposure. This implies that for exposure-dating applications we only need 997 an estimate of the effective e-folding length of production due to muons in the upper 998 few centimeters below a rock surface. Figure 16 shows an estimate of this quantity de-999 rived by using Model 1A to compute production rates due to muons at 0 and 10 g cm⁻² 1000 and deriving an effective e-folding length from these values. Note that this calculation 1001 is not possible using Model 1B because of a numerical artifact in the code of Lifton 1002 et al. (2014) that causes calculated production rates due to negative muon capture to 1003 spuriously increase with depth between 0-10 g cm^{-2} depth at some locations. 1004



Figure 16: Circles are effective e-folding lengths in upper 10 g cm⁻² for production of ¹⁰Be (red), ²⁶Al (blue), and ¹⁴C (green) by muons, calculated using Model 1A with $\alpha = 1$, calibrated to the Beacon Heights data. The solid line is the approximation given by Equation 8.

The effective e-folding length for production due to muons immediately below the surface is quite variable with elevation. This was not recognized by Balco et al. (2008), who used a single representative value without regard for elevation. As shown in Figure 16, it can be approximated by the following relationship:

$$\Lambda_{\mu,eff} = (a+bh)^{-1} \tag{8}$$

Where $\Lambda_{\mu,eff}$ is the effective e-folding length (g cm⁻²) for muon production in the uppermost 10 g cm⁻² below the surface, *h* is atmospheric pressure (hPa), *a* = 0.01036, and *b* = -9.697 × 10⁻⁶.

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To summarize this section up to this point, if we accept that Model 1A is accurate enough for typical exposure-dating applications, the approximation to Model 1A given by Equations 7 and 8 is also accurate enough for these applications. Thus, for most surface exposure dating calculations, it is not necessary to expend computation time on fully evaluating the Model 1A code. These approximations are good enough.

I now address the issue of the differences between Models 1A and 1B in scaling 1017 surface production rates due to muons, as they relate to surface exposure-dating ap-1018 plications. Model 1A assumes that the muon flux varies with atmospheric pressure, 1019 but not with geomagnetic cutoff rigidity. Model 1B assumes that the muon flux varies 1020 with both atmospheric pressure and cutoff rigidity. It also allows for variation in the 1021 muon flux with solar modulation, but this is of relatively minor importance; in apply-1022 ing Model 1B throughout this paper I assume that the solar modulation constant is the 1023 mean Holocene value according to Lifton et al. (2014). 1024



Figure 17: Surface production rate of ¹⁰Be due to muons predicted by Models 1A and 1B. Model 1A (transparent mesh) is variable only with atmospheric pressure. Model 1B (colored surface) varies with both atmospheric pressure and geomagnetic cutoff rigidity.

Figure 17 shows the difference between these scaling models, both with crosssections calibrated at Beacon Heights. Differences in predicted surface production rates due to muons are as large as 50% in some parts of the space represented in the figure. This figure also highlights an apparent numerical artifact in the Model 1B scaling method: the periodic variation in production rate with cutoff rigidity that is evident near sea level pressure in this figure is not expected from the physics of muon produc-1028 tion. This effect originates in calculations of negative muon capture production in the 1031 Model 1B code and is possibly a spurious result of differencing polynomial approxi-1032 mations for muon energy spectra used by Sato et al. (2008) and thence by Lifton et al. 1033 (2014). This apparent artifact, as well as the observation described above that Model 1034 1B appears to over-scale production rates between Beacon Heights and Cuiaba, both 1035 indicate inaccuracies in Model 1B. However, it is also clear from basic physical prin-1036 ciples that production rates due to muons should, in fact, vary with cutoff rigidity, so 1037 even though Model 1A appears to perform well to the extent that it can be tested against 1038 calibration data, it is likely oversimplified. Thus, I now ask how large an inaccuracy 1039 could be introduced into surface exposure-dating calculations if, in fact, model 1B is 1040 correct and Model 1A is not. 1041

Figure 18 shows the results of this calculation. For ²⁶Al and ¹⁰Be, not knowing 1042 whether Model 1A or Model 1B is correct has a negligible (e.g., less than 1%) effect 1043 on the total surface production rate estimate for nearly all locations. At very high cutoff 1044 rigidities (> 15 GV), which are predicted to occur rarely and infrequently by most pa-1045 leomagnetic reconstructions so are largely irrelevant for nearly all practical purposes, 1046 differences approach 2%. Figure 18 also shows that existing ¹⁰Be and ²⁶Al production 1047 rate calibration data (surface production rate calibration data relevant to calibrating 1048 spallogenic production rates, not subsurface data relevant to calibrating muon produc-1049 tion models as discussed in this paper) are subject to minimal uncertainty (< 0.5% in 1050 all cases) from this effect. Thus, not knowing which muon scaling model is preferable 1051 will not introduce any systematic bias into calibration of spallogenic production rates 1052 for ²⁶Al and ¹⁰Be using available data. However, a percent-level error might be intro-1053 duced if subsequently using this calibration to compute exposure ages at low elevation 1054 and low latitude. 1055

For ¹⁴C, on the other hand, not knowing whether Model 1A or 1B is correct introduces significant (i.e., > 5%) uncertainty in total production rate estimates at lowelevation sites. As I briefly discuss above but do not pursue further in this paper, existing ¹⁴C production rate calibration data from surface samples (Figure 18) could potentially provide some constraints on which model is correct.

It would be possible to further limit the effect of not knowing whether Model 1A or 1B are correct on surface exposure dating applications by using a smoothed approximation intermediate between the two models, that shows less variability with Rc than Model 1B but more than 1A, and also smooths out apparent polynomial artifacts in Model 1B. An example approximation with these properties for ¹⁰Be would be: 1062 1063 1064 1064 1065 1066 1066 1066 1067 1068

$$P_{\mu,10,0}(h,R_C) = \left(0.07 - \frac{0.015R_C}{20}\right) \exp\left[\frac{(1013.25 - h)}{250 + 2.5R_C}\right]$$
(9)

The maximum difference in the total surface 10 Be production rate estimate between this approximation and the predictions of either Model 1A or 1B is less than 0.2% for available calibration sites and does not exceed 0.6% for atmospheric pressure between 450-1013.25 hPa and R_C between 0-15 GV. Overall, I conclude that either the approximation to Model 1A given in Equations 7 and 8, or a smoothed intermediate model such as that given by Equation 9 with subsurface attenuation lengths estimated from Equation 8, will yield acceptable accuracy for surface exposure dating applications us-



Figure 18: Effect on total surface ¹⁰Be (top), ²⁶Al (middle), and ¹⁴C (bottom) production rate estimates (including both spallation and muons) of not knowing whether Model 1A or 1B is correct. The contoured field is the difference in predicted production rates due to muons between models as a fraction of the total surface production rate (e.g., the contour labeled '0.01' indicates that the difference between models is 1% of the total production rate). Circles are locations of ¹⁰Be and ²⁶Al calibration data from the ICE-D production rate calibration database (calibration.ice-d.org) and ¹⁴C calibration data from Borchers et al. (2016).

ing ²⁶Al and ¹⁰Be while minimizing computation time. Basically, if muons account 1073 for 2% of surface production, one should not have to devote 98% of computation time 1074 to them in surface exposure dating applications. This analysis shows that in nearly 1075 all practical applications of exposure-dating, a highly simplified approximation is ade-1076 quate. Although this conclusion also holds for ¹⁴C measurements at most locations, it 1077 may not for some measurements at low elevation and low latitude. 1078

10. Needed accuracy/precision for erosion rate estimates.

Estimating erosion rates for either single locations or entire catchments involves 1080 solving Equation 1 at the surface (z = 0) for ϵ . If we assume that production is only by 1081 fast neutron spallation and take $N_{0,i}$ to be the surface concentration and $P_{0,i}$ the surface production rate of nuclide *i*, this reduces to: 1083

$$N_{i,0} = \frac{P_i}{\lambda_i + \frac{\epsilon}{\Lambda_{xp}}} \tag{10}$$

This is explicit in ϵ so can be directly solved to yield:

$$\epsilon = \frac{\Lambda_{sp}}{N_{0,i}} \left(P_i - N_{0,i} \lambda_i \right) \tag{11}$$

This formula has been commonly used to relate nuclide concentrations in rock 1085 surfaces or stream sediment to erosion rates (e.g., Lal, 1991). If we include production 1086 by muons as well as spallation, we have: 1087

$$N_{0,i} = \frac{P_{sp,i}}{\lambda_i + \frac{\epsilon}{\Lambda_{sp}}} + \int_0^\infty P_{\mu,i}(\epsilon t) e^{-\lambda_i t} dt$$
(12)

Even if we approximate the depth dependence of production rates due to muons 1088 by an exponential function, this is still implicit in ϵ , so requires an implicit solution 1089 method. If one wants to use Model 1A or 1B to compute the right-hand term, repeated 1090 numerical integrations are necessary to evaluate the integral in the second term in the 1091 equation, and in addition the integral will need to be evaluated multiple times during 1092 the implicit solution scheme. Thus, if we accept either Model 1A or 1B as the most 1093 accurate calculation method for production rates due to muons, then solving Equation 1094 12 for the erosion rate is computationally very time-consuming. 1095

Importantly, however, in computing erosion rates from surface nuclide concentra-1096 tions by solving Equation 12, it is never necessary to accurately know the production 1097 rate due to muons at any particular depth; it is only necessary to know the integrated 1098 nuclide concentration in the right-hand term. This suggests that the muon production 1099 rate calculation could potentially be highly simplified without loss of accuracy in esti-1100 mating the erosion rate. Even though an exponential approximation for $P_{\mu,i}(z)$ would 1101 still not allow an explicit solution, it would make the implicit solution much faster. To 1102 do this, observe that there exists some value for an effective e-folding length $\Lambda_{\mu,eff,i}$ 1103 for which the second term in Equation 12 can be represented by a single exponential 1104 approximation, that is: 1105

1082

1084

$$\int_{0}^{\infty} P_{\mu,i}(\epsilon t) e^{-\lambda_{i}t} dt = \frac{P_{\mu,0,i}}{\lambda_{i} + \frac{\epsilon}{\Lambda_{\mu,eff,i}}}$$
(13)

Here $P_{\mu,0,i}$ is the surface production rate of nuclide *i* due to muons, which we can estimate from the simplified approximation to Model 1A given by Equation 7. If we evaluate the integral on the left side of this equation using the Model 1A or 1B code, we can solve the equation for $\Lambda_{\mu,eff,i}$. As described earlier in previous discussion of exponential approximations, $\Lambda_{\mu,eff,i}$ varies with both the atmospheric pressure and the erosion rate, and in the case of Model 1B it also varies with geomagnetic cutoff rigidity.



Figure 19: Variation in $\Lambda_{\mu,eff,10}$ with atmospheric pressure and erosion rate, calculated using Model 1A with $\alpha = 1$.

Figure 19 shows the variation in $\Lambda_{\mu,eff,10}$ (that is, appropriate for ¹⁰Be) for Model 1A, over the range of atmospheric pressure and erosion rate that is likely to be encountered in practice, on a 20 x 20 grid that is linearly spaced in atmospheric pressure and logarithmically spaced in erosion rate. Having computed $\Lambda_{\mu,eff,10}$ on a grid spanning a range of pressure and erosion rate, one can simplify Equation 12 as follows:

$$N_{0,i} = \frac{P_{sp,i}}{\lambda_i + \frac{\epsilon}{\Lambda_{sp}}} + \frac{P_{\mu,i}}{\lambda_i + \frac{\epsilon}{\Lambda_{\mueffi}(h,\epsilon)}}$$
(14)

where the function $\Lambda_{\mu,eff}(h,\epsilon)$ is defined in discrete form by the precalculated grid shown in Figure 19. This is still implicit, but if one uses Equation 7 above to compute $P_{\mu,i}$ at the site, and then evaluates the function $\Lambda_{\mu,eff,i}(h,\epsilon)$ by interpolation from the precalculated grid shown in Figure 19, solving it is computationally trivial. Tests of this procedure against solving Equation 12 by numerical integration using the complete Model 1A code shows that results of the simplified procedure differ from results calculated using the complete code by less than 0.25% for the entire range of atmospheric pressure and erosion rate shown.

A similar procedure could be developed for Model 1B, except that one would need 1125 to determine the dependence of $\Lambda_{\mu, eff, i}(h, \epsilon)$ on geomagnetic cutoff rigidity as well as 1126 atmospheric pressure and erosion rate, thus requiring interpolation on a 3-dimensional 1127 instead of 2-dimensional grid. To determine if this is necessary, I did an experiment 1128 in which I calculated erosion rates at an array of sites spanning a range of atmospheric 1129 pressure, cutoff rigidity, and erosion rate using both (i) numerical integration of the full 1130 Model 1B code, and (ii) the simplified procedure based on values of $\Lambda_{u,eff}(h,\epsilon)$ com-1131 puted using the Model 1A code and shown in Figure 19. In other words, I assumed that 1132 Model 1B is correct, but solved Equation 14 using values for $\Lambda_{\mu,eff,i}(h,\epsilon)$ computed 1133 using Model 1A. Despite intentionally using internally inconsistent calculation meth-1134 ods in this way, maximum differences in resulting erosion rate estimates were still less 1135 than 5% in all cases and less than 1% in 93% of cases. This shows that uncertainty in 1136 whether Model 1A or Model 1B is correct contributes negligible uncertainty to erosion 1137 rate estimates based on ¹⁰Be concentrations. 1138

To summarize, one can use a simple exponential approximation for ¹⁰Be or ²⁶Al 1139 production by muons for erosion rate calculations without loss of accuracy, as long 1140 as the calculation method captures the fact that the effective e-folding length applica-1141 ble to this calculation varies with location and erosion rate. In the paragraphs above 1142 I have proposed a method to do this by computing $\Lambda_{\mu,eff}(h,\epsilon)$ as a function of h and 1143 ϵ by interpolation from a coarse grid of precalculated values, which is computation-1144 ally extremely simple and maintains accuracy at the < 1% level compared to using the 1145 complete Model 1A code. Note that although I have represented this idea by defining 1146 an effective e-folding length $\Lambda_{\mu, eff, i}$ that varies with atmospheric pressure and erosion 1147 rate, in an actual computational implementation it would be faster and simpler to pre-1148 compute and store values for the entire second term in Equation 12 on a grid in erosion 1149 rate and atmospheric pressure that could be used for interpolation, rather than taking 1150 the redundant steps of representing these results as values of $\Lambda_{\mu,eff,i}$ and then using 1151 them to recalculate the value of the integral within the implicit solver. 1152

A final note in this section is that in practice the precision of the numerical method 1153 of computing an erosion rate from a surface nuclide concentration is not an important 1154 limit on the absolute accuracy of the erosion rate estimate. A much more important 1155 limit on the accuracy of the erosion rate estimate is that the calculation is based on the 1156 assumption that the muon-produced nuclide inventory has reached equilibrium with a 1157 steady erosion rate. For relatively long-half-life nuclides like ¹⁰Be and ²⁶Al, in nearly 1158 all geological situations that can be envisioned, this is unlikely to be the case. Thus, no 1159 matter what the precision of the numerical solution method, whether or not the erosion 1160 rate estimate accurately represents the true erosion rate is much more sensitive to the 1161 geological assumptions necessary to do the calculation at all. 1162

11. Conclusions

Models for production of ²⁶Al, ¹⁰Be, and ¹⁴C in quartz by cosmic-ray muons that ¹¹⁶⁴ are based on downward propagation of the surface muon energy spectrum using the ¹¹⁶⁵

method described by Heisinger can, in general, successfully match existing subsurface calibration data. However, using these calibration data as a quantitative test for the accuracy of methods for geographic scaling of the surface muon flux is not conclusive, because erosion rates at most available calibration sites are too high to effectively decouple production rate estimates from the steady-state erosion assumption.

A possible exception is that a comparison between the two lowest-erosion-rate sites 117 at Beacon Heights (high elevation, high latitude) and Cuiaba (low elevation, low lati-1172 tude) shows that the implementation of Heisinger's method in Balco et al. (2008) (re-1173 ferred to here as Model 1A), appears to perform better than the similar implementa-1174 tion in Lifton et al. (2014) that uses more complex geographic scaling (here, Model 1175 1B). Physical arguments indicate that geographic scaling in Model 1B, which includes 1176 variation in the muon flux with magnetic cutoff rigidity, should be more accurate for 1177 geographic scaling of muon flux than Model 1A, which does not include such vari-1178 ation. However, Model 1B overestimates the difference in production rates between 1179 Beacon Heights and Cuiaba. In addition, the Model 1B code of Lifton et al. (2014) 1180 predicts some physically unexpected and likely spurious production rate variations that 1181 may originate from approximations used in the code. Thus, despite the expectation that 1182 Model 1B should perform better than Model 1A, available data indicate that it does not. 1183 This conclusion, however, is weak and could be better evaluated by generating new cal-1184 ibration data from sites with erosion rates on the order 0.1 m Myr⁻¹ or less, at lower 1185 elevation and lower latitude than Beacon Heights. Preferably these data would be from 1186 a borehole in a bedrock surface where sample depths are accurately known and not 1187 subject to the uncertainties that arise from opportunistic sampling at mine excavations. 1188

Burial-dating and depth-profile dating applications require precise estimates of pro-1189 duction rates due to muons at specific depths, so in nearly all cases for these applica-1190 tions it is necessary to use one of the models based on downward propagation of the 1191 muon energy spectrum, that is, either Model 1A or 1B. An upper limit on the absolute 1192 uncertainty in subsurface production rates for ²⁶Al and ¹⁰Be computed using Model 1193 1A, to the extent it can be determined from available calibration data, is 25% globally, 1194 but it is likely that actual uncertainties are closer to 10%, similar to scaling uncertainties 1195 for spallogenic production. For burial-dating and depth-profile applications, simplified 1196 exponential approximations to subsurface production rates can be fit to a limited depth 1197 range at a specific site in order to speed up production rate calculations relative to the 1198 complete Model 1A or 1B codes. However, global scaling models based on a simple 1199 exponential approximation to calibration data are oversimplified for this purpose. 1200

For surface exposure dating applications using ¹⁰Be and ²⁶Al in quartz, the muonproduced nuclide inventory is typically a small fraction of the total measured nuclide concentration in surface samples. Thus, a highly simplified scaling method that represents the surface production rate due to muons by a simple exponential function of atmospheric pressure can be used without significant loss of accuracy in relation to more complex muon scaling models. With some exceptions, this is likely also the case for exposure-dating with ¹⁴C in quartz.

For estimating steady erosion rates from surface ¹⁰Be or ²⁶Al concentrations, it is ¹²⁰⁸ also possible to develop a simple and computationally trivial scaling method, based on ¹²¹⁰ an effective attenuation length for muon production that is variable with location and ¹²¹⁰ erosion rate, that maintains the numerical precision of the calculation in relation to the ¹²¹¹

computationally time-consuming Heisinger method.	1212
12. Computer code	1213
All calculations in this paper were done using MATLAB (R2015b). Many also re-	1214
quire the MATLAB Optimization Toolbox. MATLAB code to perform all calculations	1215
and generate all figures is available at the following address:	1216
http://hess.ess.washington.edu/repository/muons2016/	1217
13. Acknowledgements.	1218
This work was supported by the Ann and Gordon Getty Foundation. I thank Régis	1219
Braucher and Martin Lupker for helpful and comprehensive reviews.	1220

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Tables

	¹⁰ Be							²⁶ AI						
Model	f*	σ ₀ (μb)	σ ₁₉₀ (μb)	Po	Λ_{μ}	Scatter (%) All data	Scatter (%) > 1000 g cm ⁻²	f*	σ ₀ (μb)	σ ₁₉₀ (μb)	Ρ	Λ_{μ}	Scatter (%) All data	Scatter (%) > 1000 g cm ⁻²
Model 1A (α = 0.75) - see Figure 4 Model 1A (α = 1) - see Figure 5	0.00157 0.00191	0.739 0.280	37.8 53.2			5.2 5.0	5.2 4.7	0.0118 0.0133	10.19 3.89	521 739			10.8 10.5	13.3 12.9
Model 1B (α = 1) - see Figure 6	0.00192	0.237	45.0			5.0	4.8	0.0131	3.26	619			10.5	13.0
Borchers et al. (2015) (similar to Model 1B)	0.00188	0.252	47.9					0.0121	4.03	766				
Laboratory measurements by Heisinger et al. 1	0.0039 ± 0.0003		94 ± 13					0.022 ± 0.002		1410 ± 170				
Model 2 - see Figure 7				0.084	2501		22.9				0.761	2417		23.7

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Table 1. Best-fitting values of muon production parameters for various production models applied to the Beacon Heights data.

 $^{\rm 1N}$ ormalized to the 07KNSTD and KNSTD standards for $^{\rm 10}{\rm Be}$ and $^{\rm 26}{\rm Al}$ respectively.

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Table 2.

Core	Latitude	Longitude	Elevation	Atmospheric pressure (hPa)	Rc (GV) (see text)	Production ra spallation (at (Borchers et ¹⁰ Be	ates due to coms/g/yr) t al., 2016) ²⁶ Al	Mean meası uncertainty > 1 (%) ¹⁰ Be	urement 1000 g/cm2 1 ²⁶ Al	Erosion rate est from spallogeni E (sp) (¹⁰ Be) (m/Myr)	timate ic inventory E (sp) (²⁶ AI) (m/Myr)	Results of mod ¹⁰ Be, Model 1A E (muons) (m/Myr)	lel fitting Scatter (%)	^o Be, Model 1B E (muons) (m/Myr)	2 Scatter (%)	^s Al, Model 1A E (muons) (m/Myr)	2 Scatter (%)	⁶ Al, Model 1B E (muons) (m/Myr)	Scatter (%)
La Ciotat	43.179	5.576	310	979.3	4.6	5.04	35.14	10.2	21	21.2 ± 2.2	18.1 ± 2.2	33.4	2.1	28.0	2.4	33.3	27.5	26.8	31.5
Leymon High	42.064	7.013	1246	874.6	ß	10.75	74.91	9.4	20.4	17 ± 1.8	17 ± 2	12.1	13.8	11.5	14.8	14.4	18.6	13.7	19.8
Leymon Low	42.065	7.014	1277	871.3	Ŋ	11.02	76.72	8.5	18.9	7.7 ± 0.8	13.5 ± 1.6	8.61	9.9	8.0	10.2	10.0	23.5	9.4	24.5

Notes: Motes: Sector rates are converted to m/Myr assuming a rock density of 2.7 g/cm2 Estimates of atmospheric pressure and magnetic cutoff rigidity (Rc) are described in the text. Production rate estimates use St' scaling scheme and calibration data of Borchers et al. (2016). Uncertainties on erosion rate estimates from spallogenic inventory include uncertainty in production rate and nuclide concentrations.